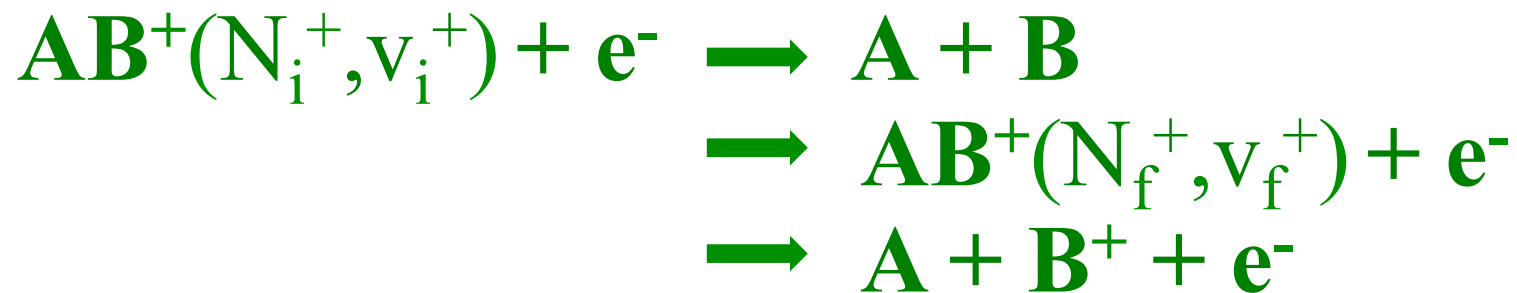




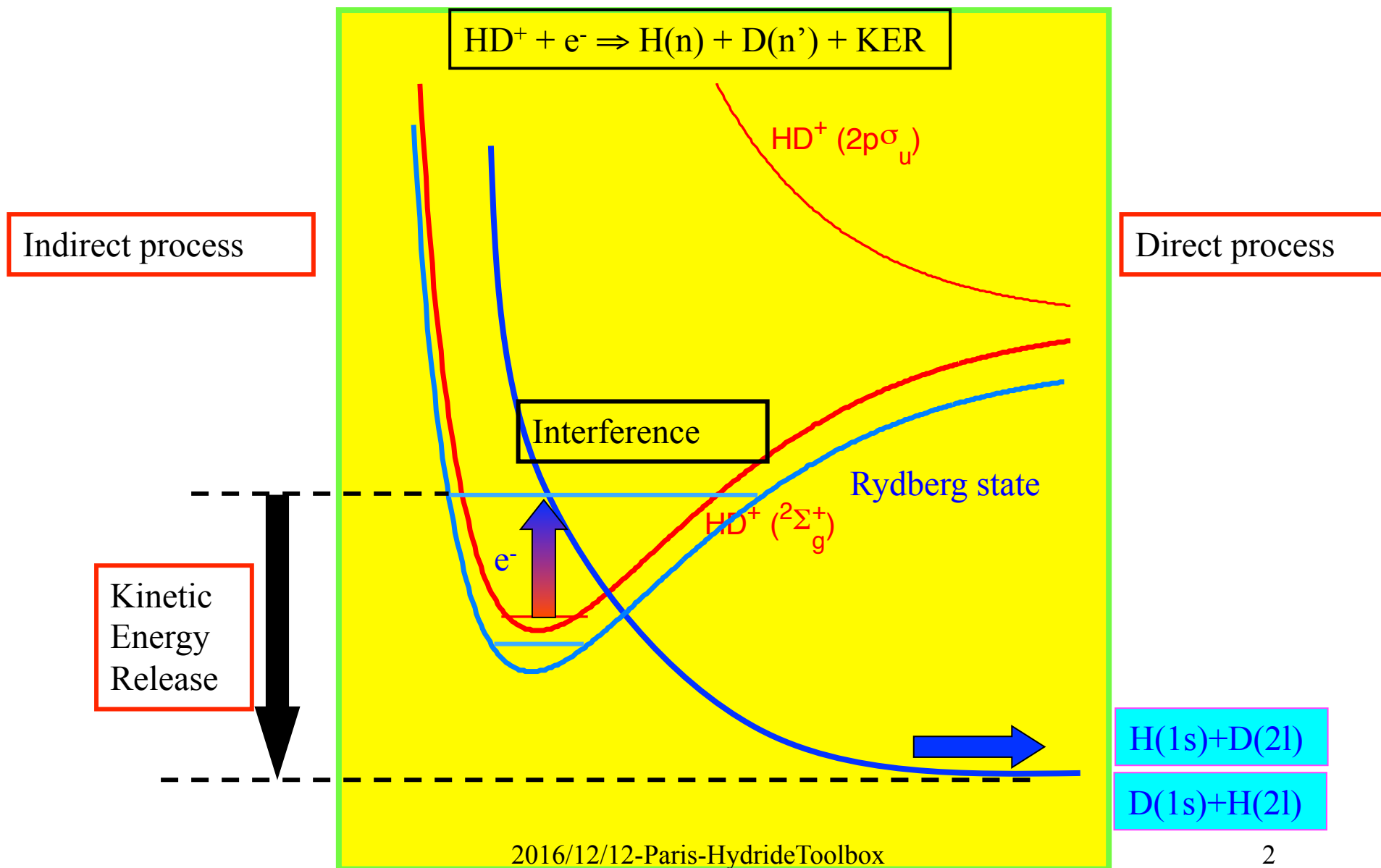
Recombination, excitation and dissociation of hydride molecular cations in low energy electron collisions

**J. Zs Mezei¹⁻⁴, F. Colboc¹, D. O. Kashinsky⁵, D. A. Little⁶, M. D. Epée Epée⁷,
S. Niyonzima⁸, N. Pop⁹, C. M. Coppola¹⁰, K. Chakrabarti¹¹, O. Motapon⁷,
A. Bultel¹², K. Hassouni², D. Talbi¹³, A. P. Hickman¹⁴, A. Faure¹⁵, J. Tennyson⁶
and I. F. Schneider^{1,3}**

¹LOMC, Univ. du Havre, ²LSPM, Univ. Paris 13, ³Lab. Aimé Cotton, Univ. Paris-Sud,
⁴Inst. of Nuclear Res. of the Hungarian Acad. of Sciences, Debrecen, ⁵US Military Academy, NY,
⁶University College London, ⁷Univ. of Douala, ⁸Univ. of Burundi, ⁹Politehnica Univ. Timisoara,
¹⁰INAF Bari, ¹¹Scottish Church College, Calcutta, ¹²CORIA, Univ. de Rouen,
¹³LUPM, Univ. de Montpellier, ¹⁴Lehigh University, PA, ¹⁵IPAG, Univ. de Grenoble.



Electron-cold molecular ion reaction: Dissociative Recombination



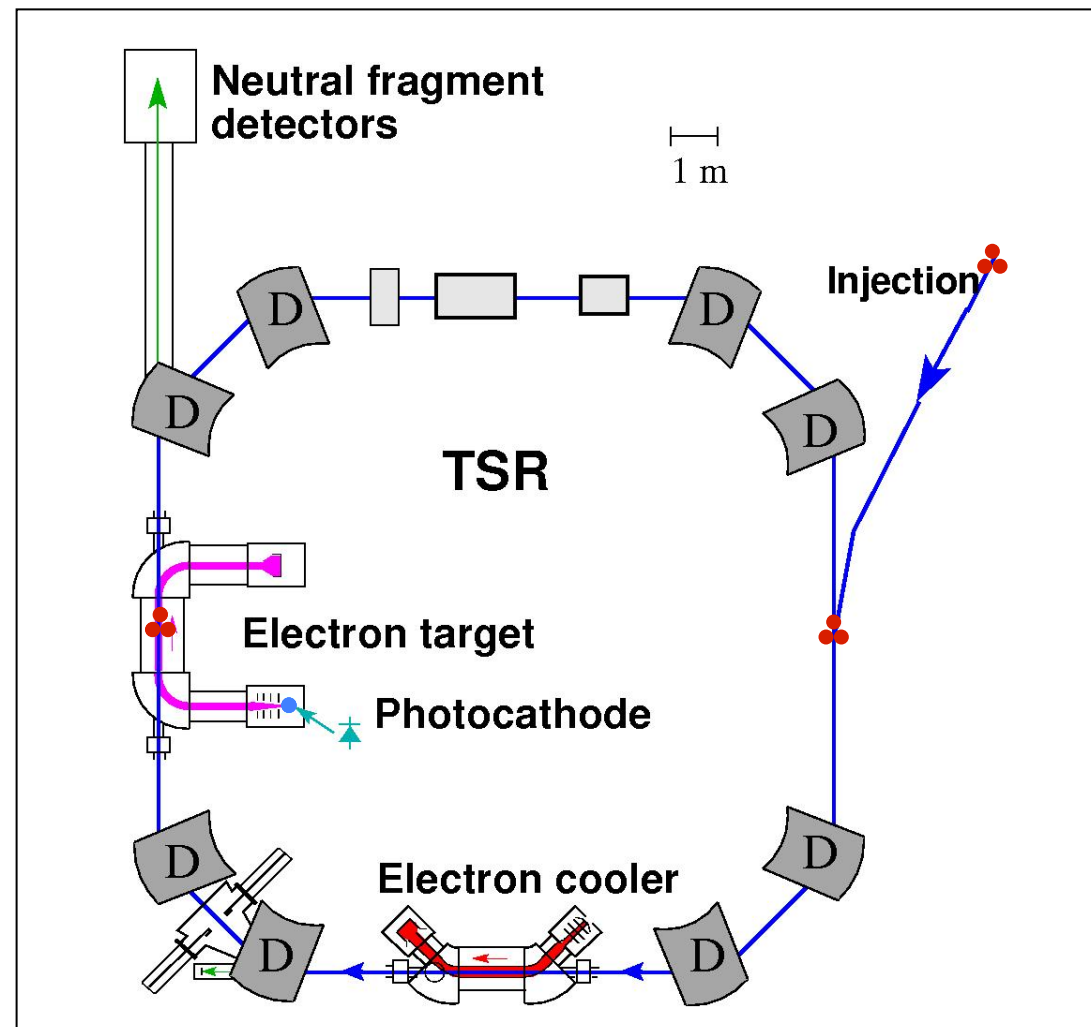
Storage ring measurements: concept

Advantages

- radiative relaxation of internal excitations (rotation, vibration)
- direct measurement
- 100% detection efficiency
- high resolution

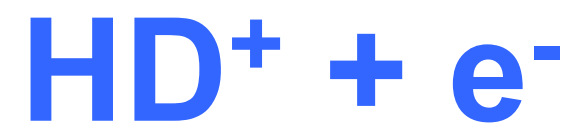
• Vibrations cool ✓

• Rotations stay forever ✗



Kreckel et al., PRA A 66, 052509 (2002),
Kreckel et al., New J. Phys. 6, 151 (2004)

Benchmarking:
a NOT-SO-Hydride system,
H₂⁺ (or isotopologues) + e⁻ :
maximum ACCURACY



Rotational transitions induced by collisions of HD⁺ ions with low-energy electrons

O. Motapon,^{1,2} N. Pop,³ F. Argoubi,⁴ J. Zs Mezei,^{2,5,6} M. D. Epee Epee,¹ A. Faure,⁷ M. Telmini,⁴
 J. Tennyson,⁸ and I. F. Schneider^{2,5}

HD⁺ + e⁻ → HD* → H + D
 +
 Average
 on
 ROTATIONAL
 distribution,
 T_{rot} ... + ...

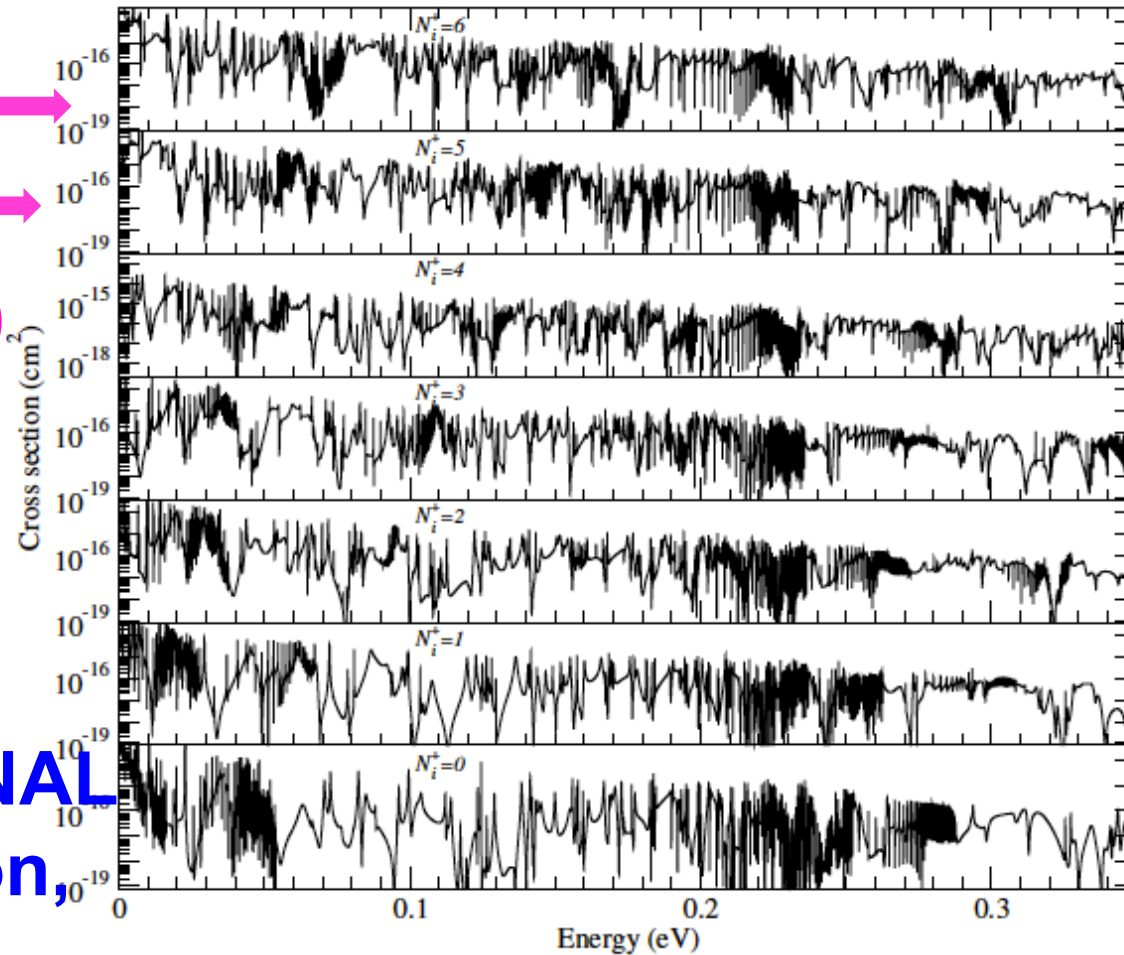
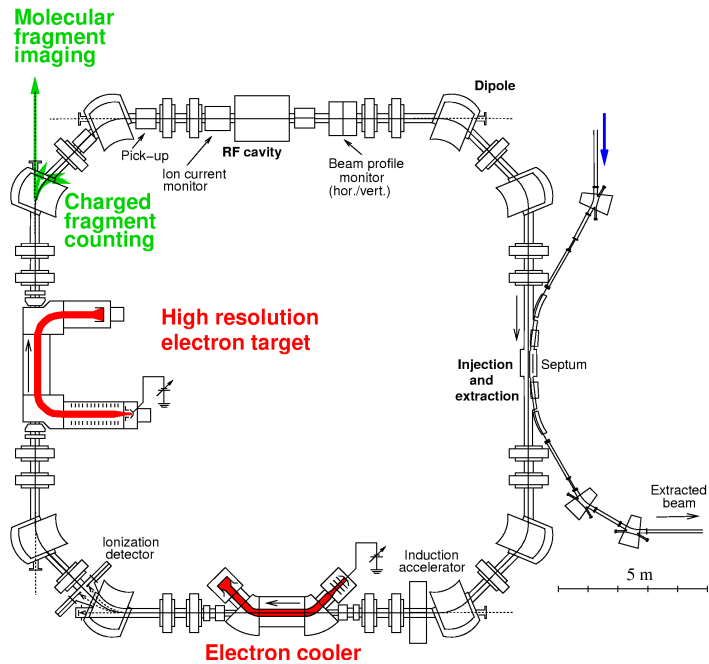


Figure 3. DR cross sections of HD⁺ initially in one of its lowest rotational level N_i^+ (vibrational ground state).



...+
**Convolution
 with
 ANISOTROPIC
 Maxwell
 Distribution:**

$$\alpha = \langle v\sigma \rangle = \iiint \sigma(v) v f(v_d, \mathbf{v}) d\mathbf{v} \quad (1)$$

$$f(v_d, \mathbf{v}) = \frac{m}{2\pi k T_{e\perp}} \exp\left(-\frac{mv_{\perp}^2}{2kT_{e\perp}}\right) \sqrt{\frac{m}{2\pi k T_{e\parallel}}} \exp\left(-\frac{m(v_{\parallel} - v_d)^2}{2kT_{e\parallel}}\right) \quad (2)$$

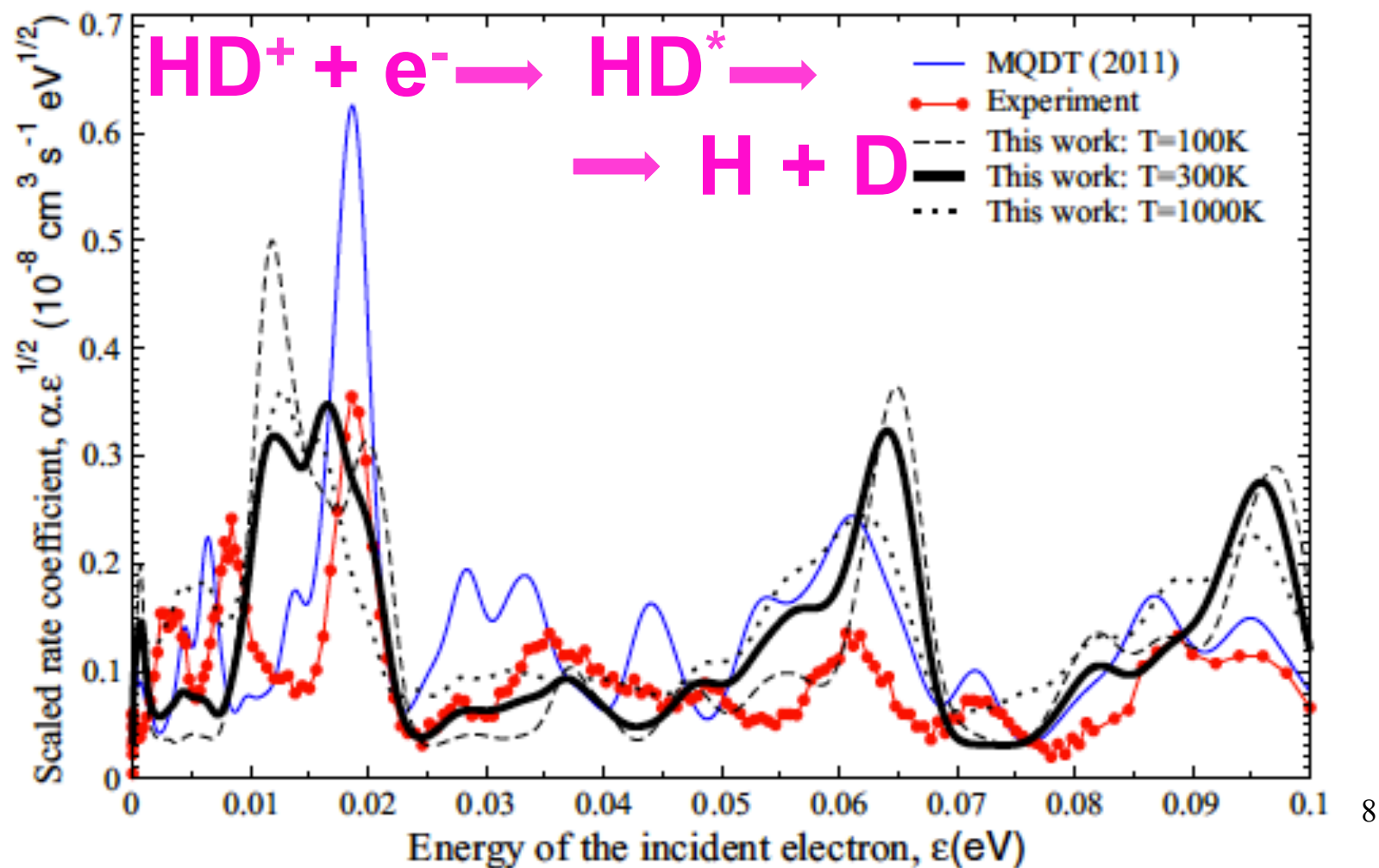
Best parameters:

$$T_{\text{long}} = 20 \mu\text{eV} = 0,23 \text{ K} = 0,16 \text{ cm}^{-1}$$

$$T_{\text{trans}} = 500 \mu\text{eV} = 5,80 \text{ K} = 4,03 \text{ cm}^{-1}$$

Rotational transitions induced by collisions of HD^+ ions with low-energy electrons

O. Motapon,^{1,2} N. Pop,³ F. Argoubi,⁴ J. Zs Mezei,^{2,5,6} M. D. Epee Epee,¹ A. Faure,⁷ M. Telmini,⁴
 J. Tennyson,⁸ and I. F. Schneider^{2,5}



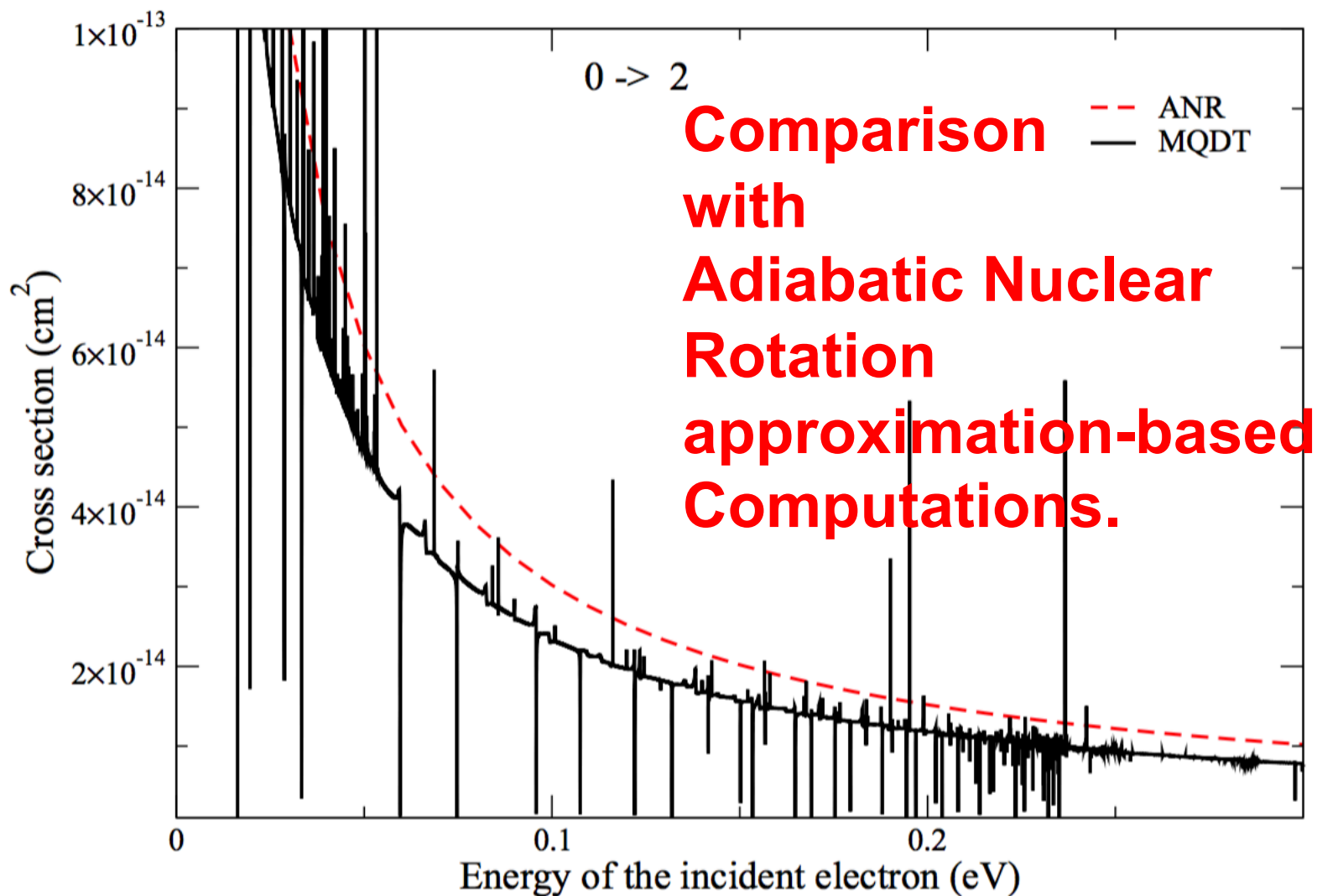


FIG. 2. (Color online) Zoom in the linear scale of the cross section for the transition $0 \rightarrow 2$, in the ground vibrational state of $\text{HD}^+(X^2\Sigma_g^+)$. Black solid curve: MQDT computations; red dashed curve: ANR approximation-based computations.

For the modelers:
Maxwell
rate-coefficients

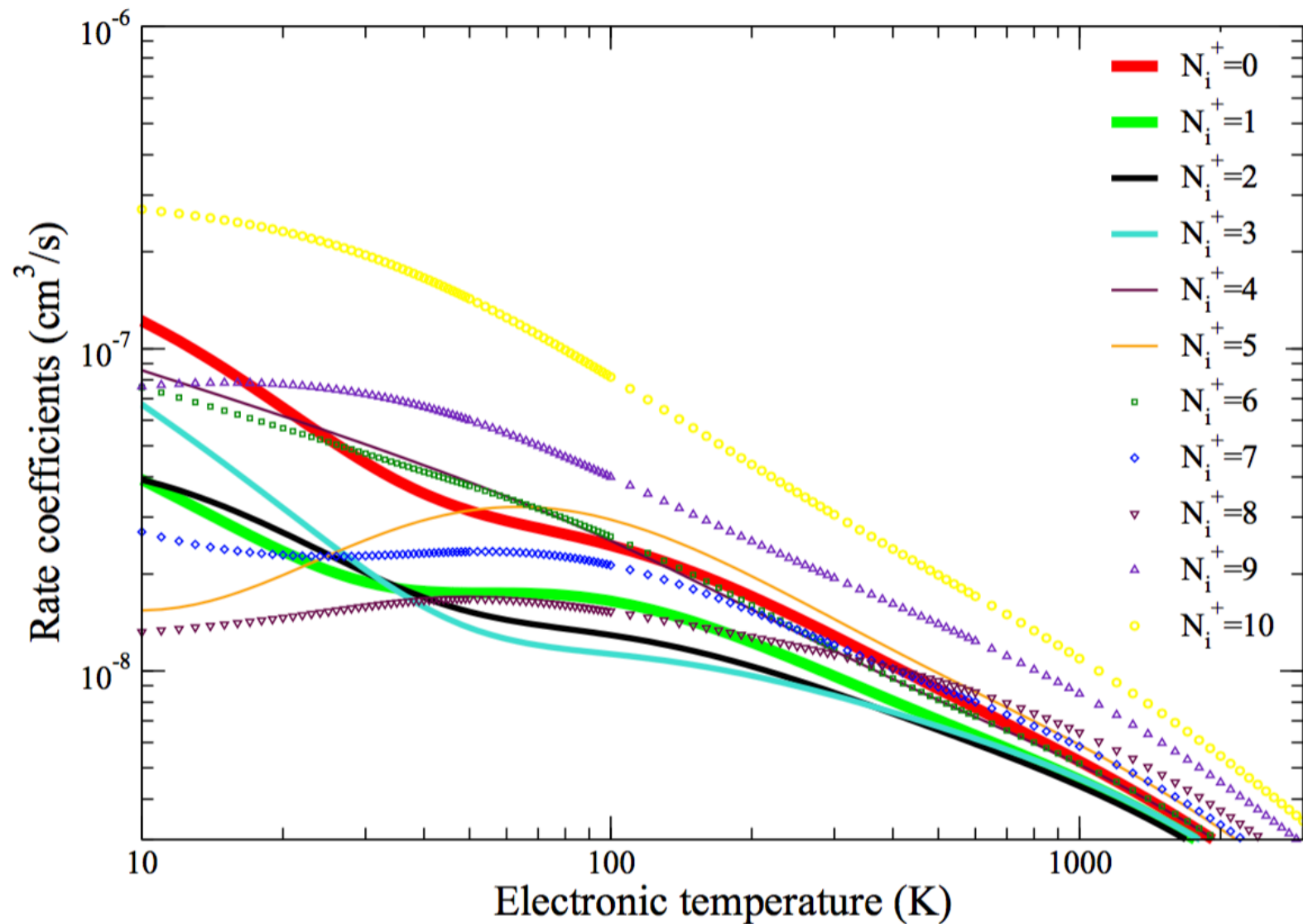


FIG. 9. (Color online) Maxwell isotropic rate coefficients for the dissociative recombination $\text{HD}^+(X^2\Sigma_g^+)$ with $v_i^+ = 0$ as a function of initial rotational level, $N_i^+ = 0$ to 10.

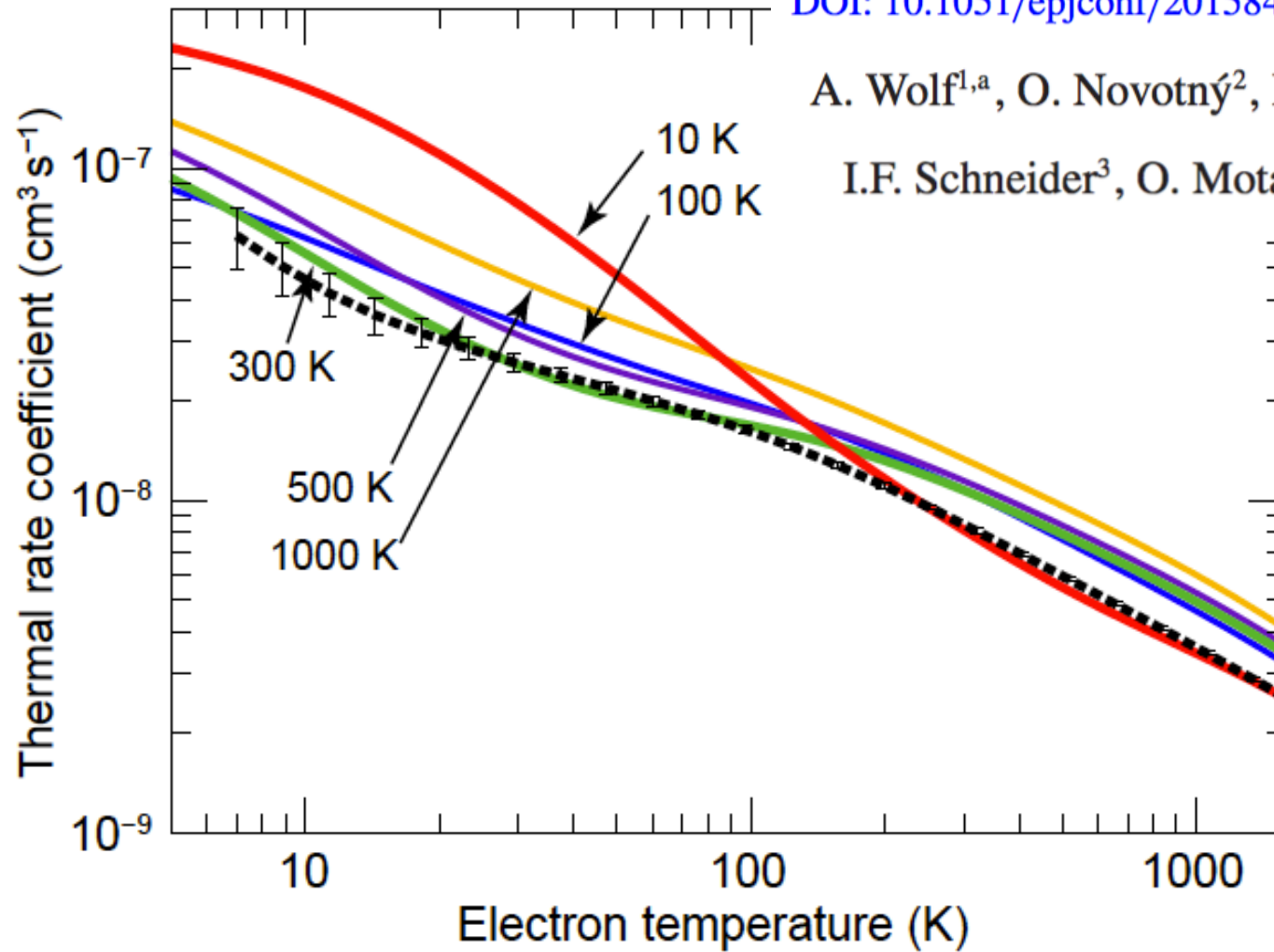
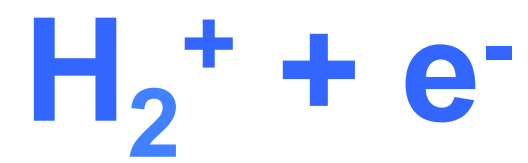
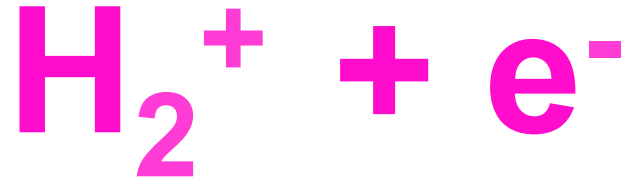
A. Wolf^{1,a}, O. Novotný², H. Buhr², C. Krantz²,I.F. Schneider³, O. Motapon⁴ and J.Zs. Mezei^{3,5}

Fig. 2. Thermal rate coefficient for HD^+ DR as a function of the electron temperature. Experiment is given by a thick dashed line; 1σ estimated experimental errors due to uncertainties in the electron beam temperatures are also indicated. Theory is obtained for the given rotational temperatures.





Reactive collisions of very low-energy electrons with H_2^+ : rotational transitions and dissociative recombination

M. D. Epée Epée,¹ J. Zs Mezei,^{2,3,4} O. Motapon,^{1,5★} N. Pop⁶ and I. F. Schneider^{2,3★}

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³*Laboratoire Aimé Cotton CNRS-UPR-3321, Université Paris-Sud, Orsay F-91405, France*

⁴*Laboratoire des Sciences des Procédés et des Matériaux, UPR 3407 CNRS and Univ. Paris 13, 99 avenue Jean-Baptiste Clément, F-93430 Villetaneuse, France*

⁵*Faculty of Science, University of Maroua, PO Box 814 Maroua, Cameroon*

⁶*Department of Physical Foundation of Engineering, University Politechnica of Timisoara, Bv Vasile Parvan No 2, 300223, Timisoara, Romania*

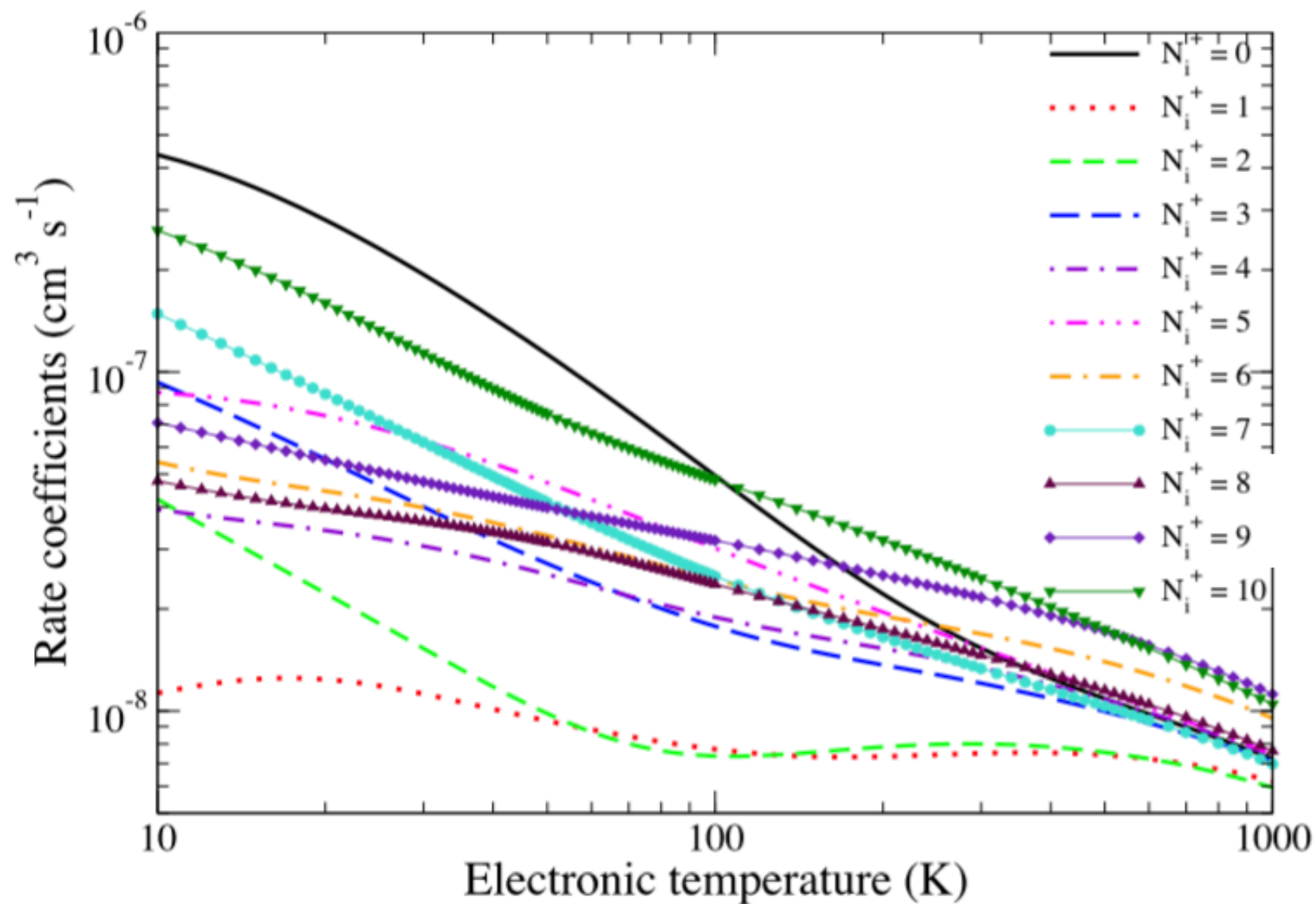


Figure 4. Maxwell rate coefficients for the DR of $\text{H}_2^+(X^2\Sigma_g^+)$ on its ground vibrational level $v_i^+ = 0$, as a function of its initial rotational level, $N_i^+ = 0$ to 10.

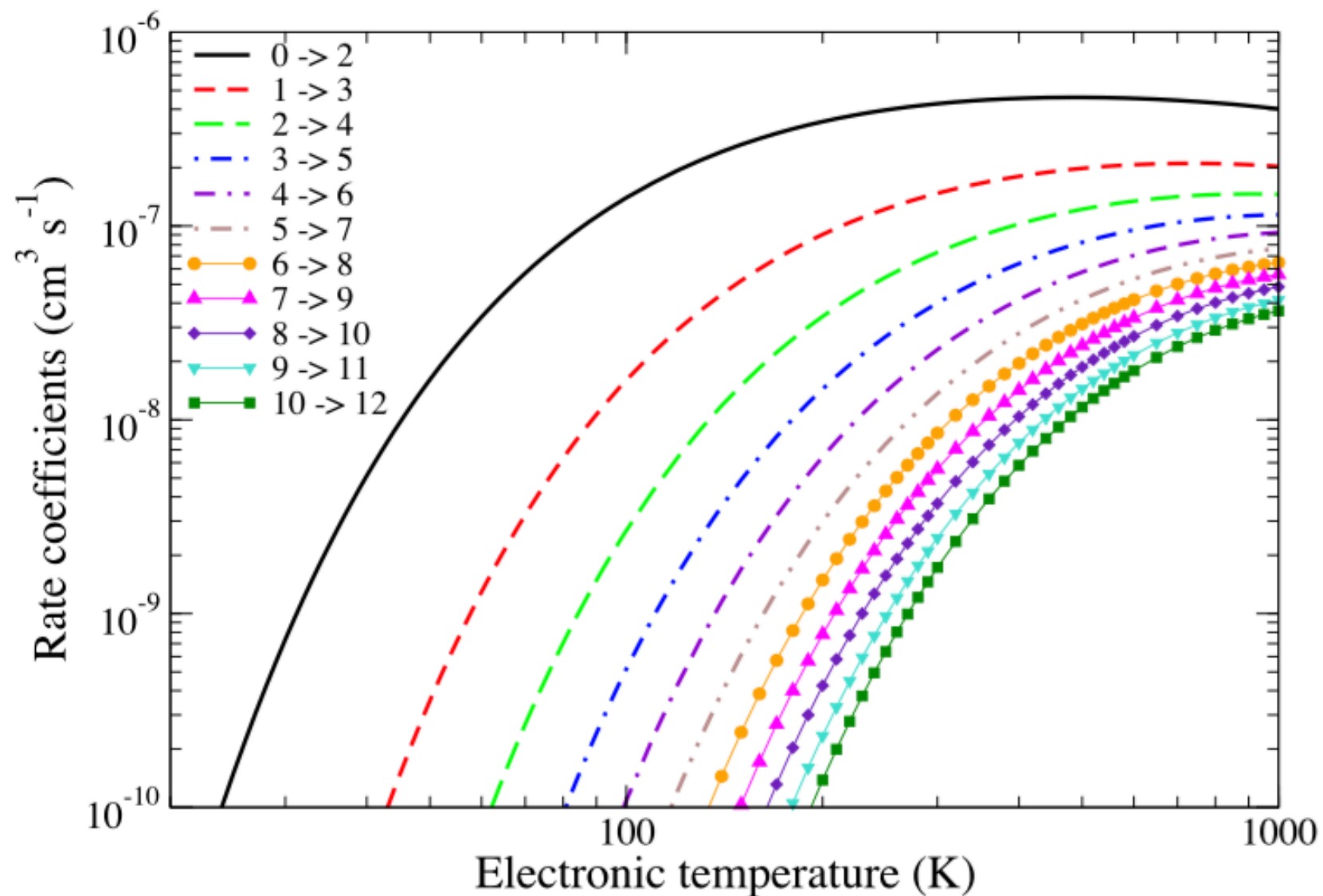
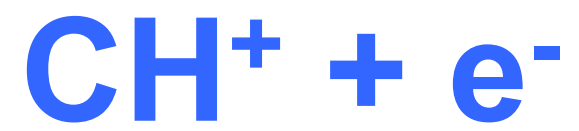
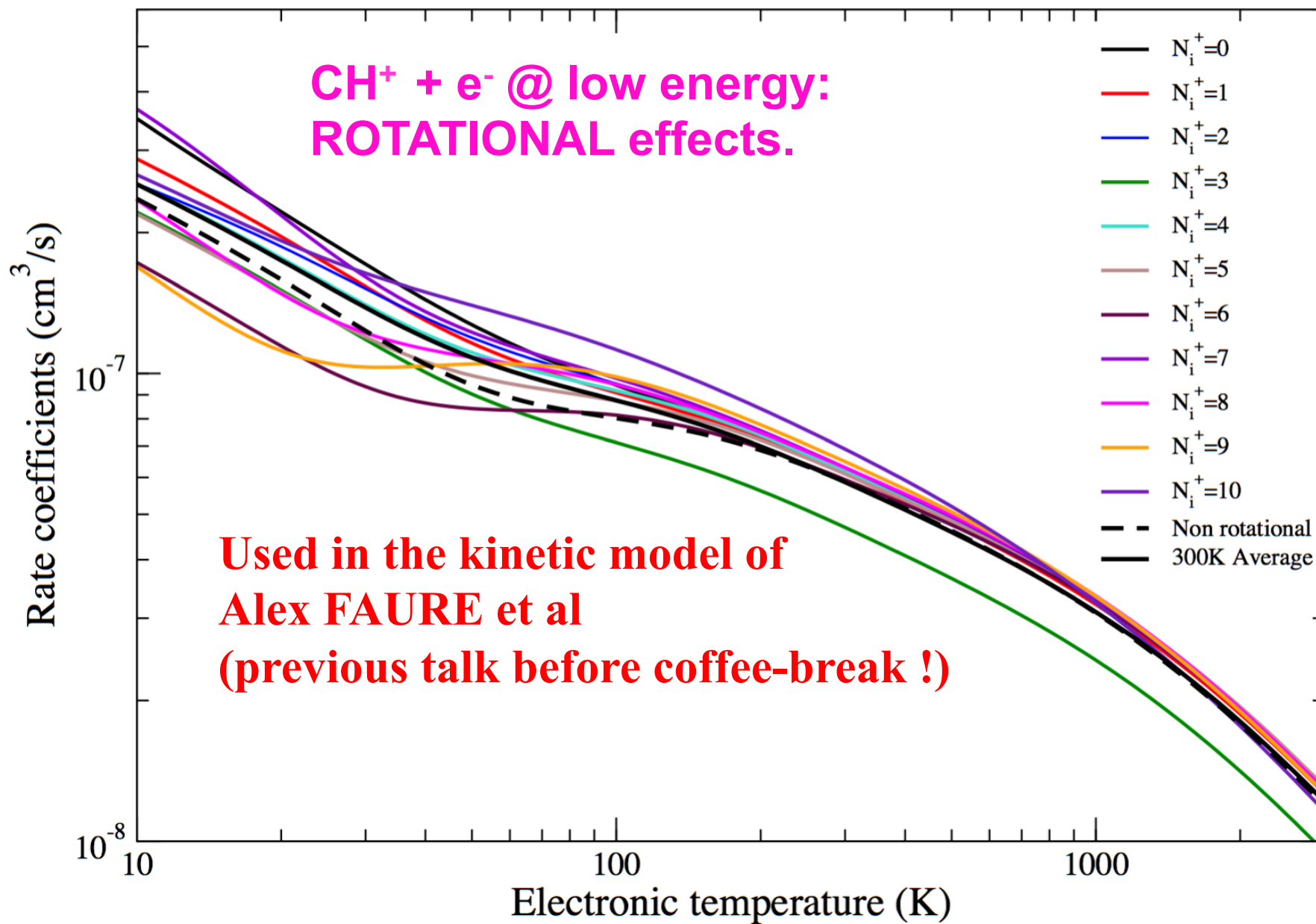
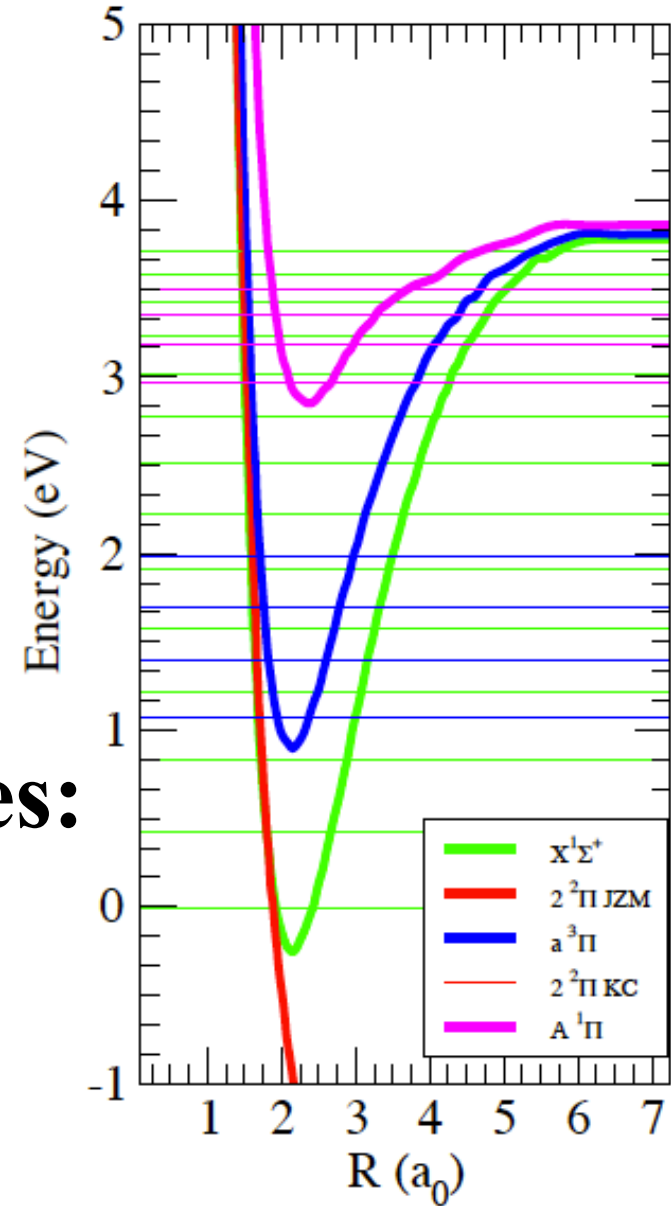
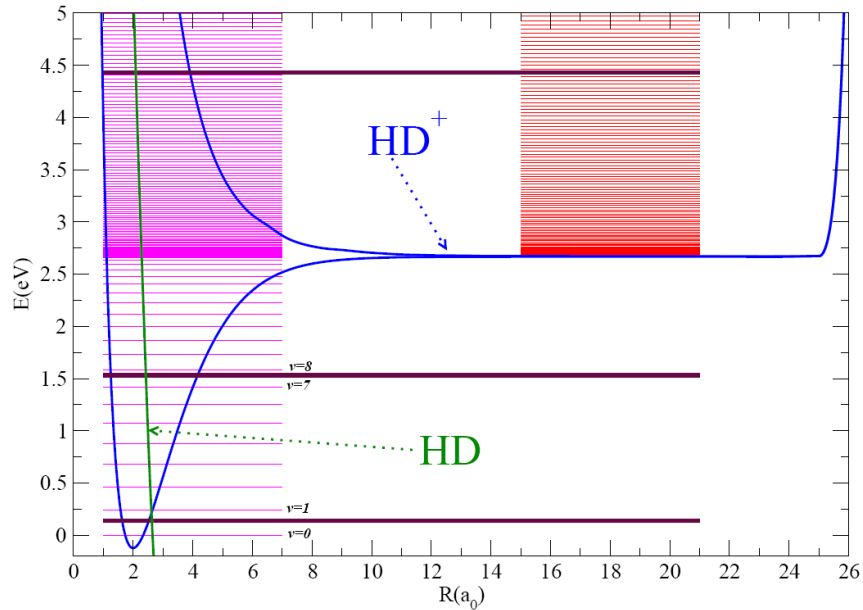


Figure 3. Maxwell rate coefficients for rotational excitation $N_i^+ \rightarrow N_i^+ + 2$, with $N_i^+ = 0$ to 10 of $\text{H}_2^+(X^2\Sigma_g^+)$ on its ground vibrational level $v_i^+ = 0$.¹⁶





CH⁺ + e⁻ @ high energy



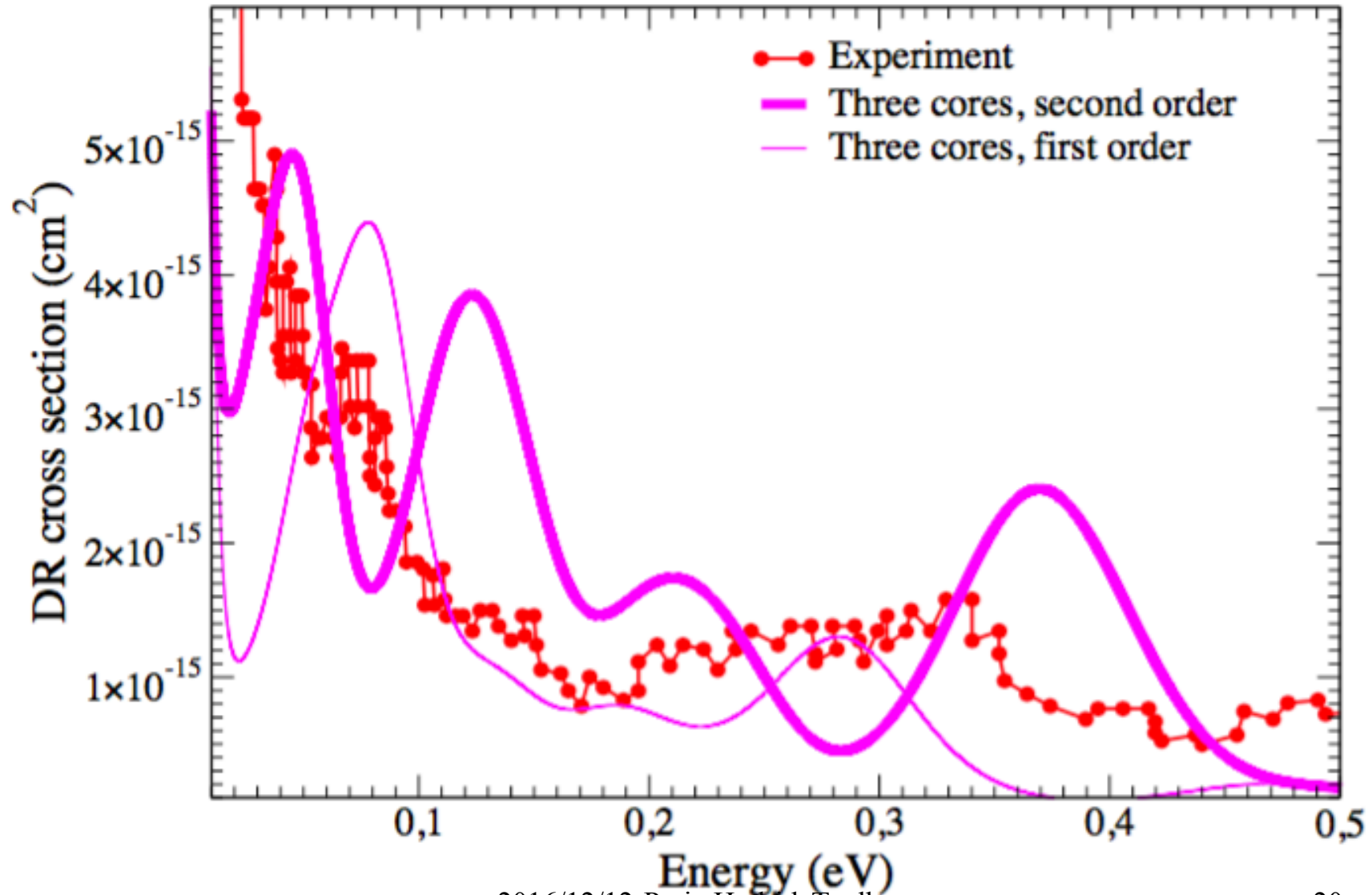
CH⁺: **BOUND** ionic cores:

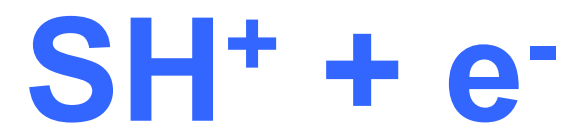
c1, c2, c3

+

electron

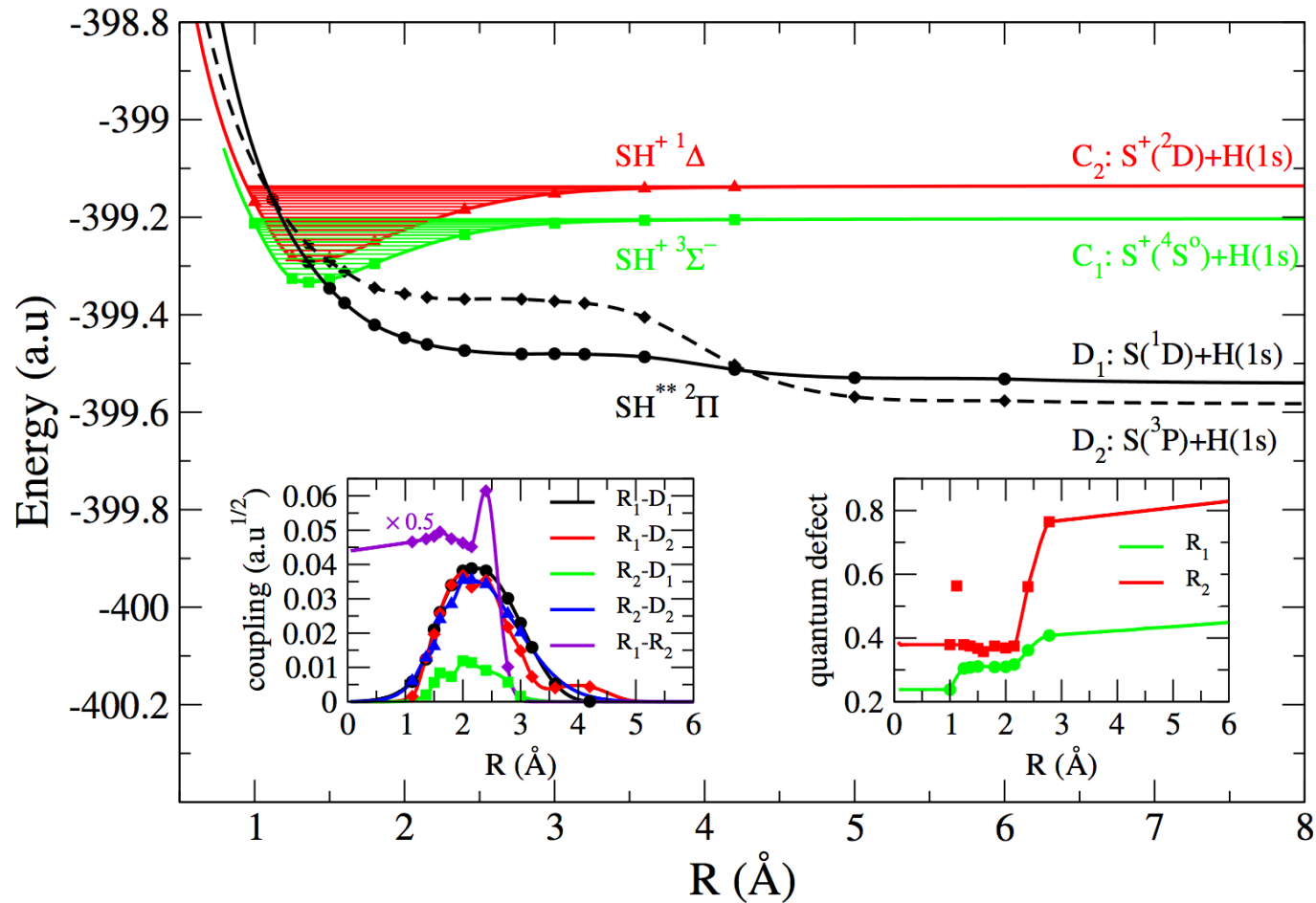
**CH⁺ + e⁻ @ high energy: rotational effects neglected,
but account of the **core-excited Rydberg states:****



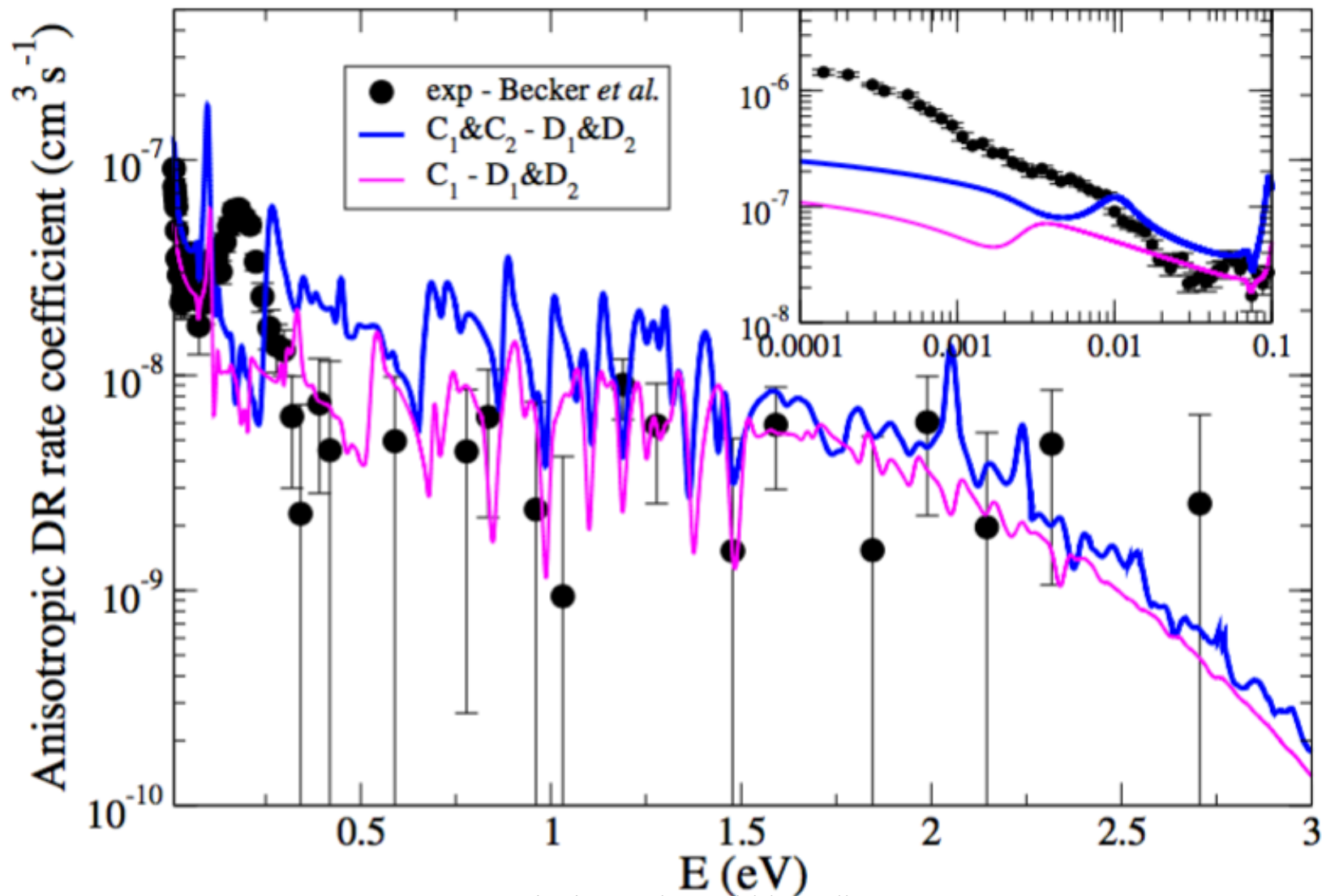


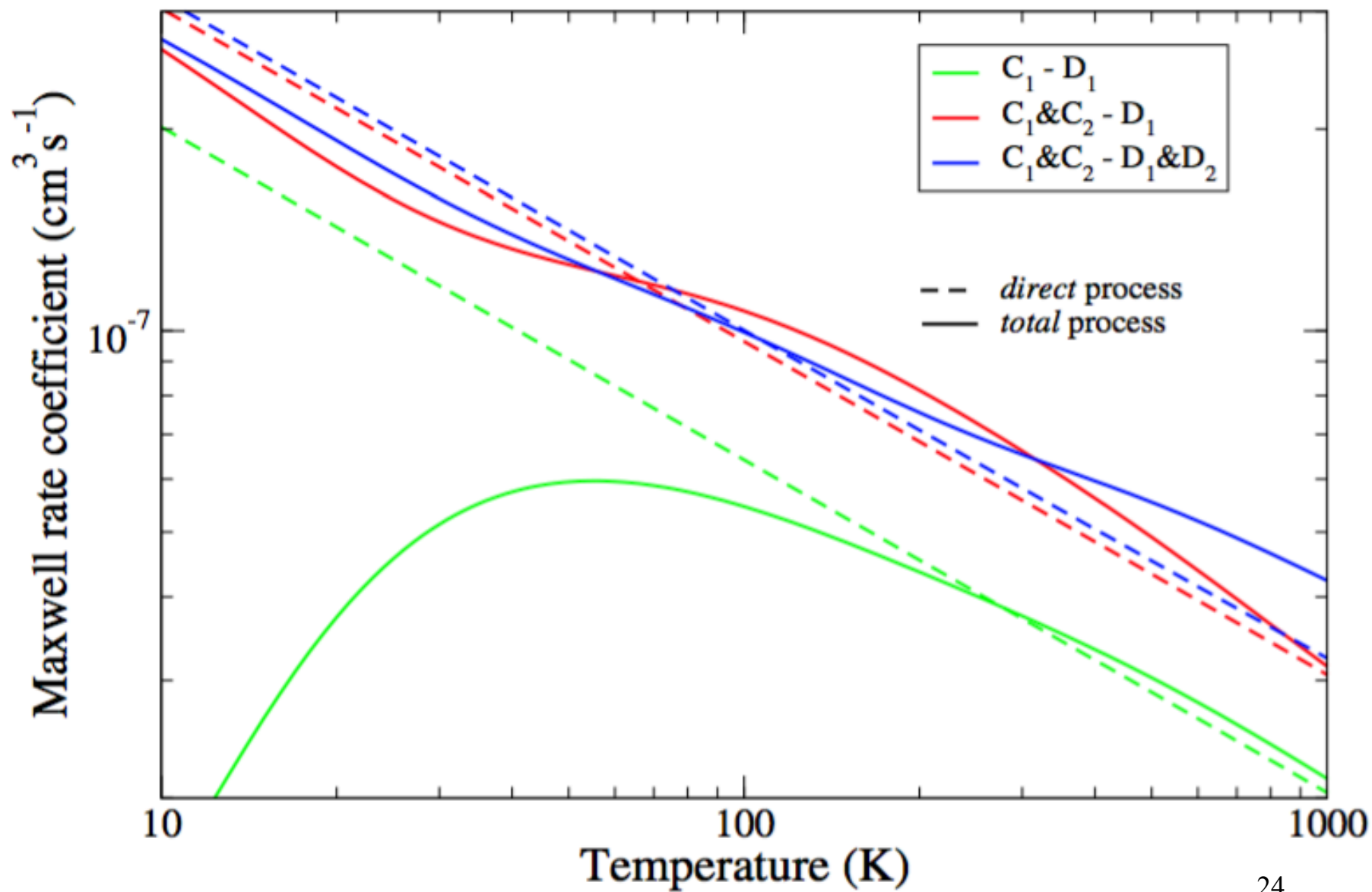
**SH⁺ + e⁻ : rotational effects neglected,
account of the core-excited Rydberg states:**

Bock-diagonalization method, Talbi et al 2015:



**SH⁺ + e⁻ : rotational effects neglected,
account of the core-excited Rydberg states:**









Beryllium monohydride (BeH): Where we are now, after 86 years of spectroscopy

Nikesh S. Dattani

Due to its simplicity, BeH is expected to be present in astronomical contexts such as exoplanetary atmospheres, cool stars, and the interstellar medium [27], but in the context of astronomy, has only been found on our Sun, in the two studies described in [28, 29].

ExoMol line lists – I. The rovibrational spectrum of BeH, MgH and CaH in the $X^2\Sigma^+$ state

Benjamin Yadin, Thomas Veness, Pierandrea Conti, Christian Hill,
Sergei N. Yurchenko and Jonathan Tennyson[★]

Department of Physics and Astronomy, University College London, Gower Street, WC1E 6BT London

BeH is one of the simplest heteronuclear diatomic molecules, and hence it is a strong contender for being observed in contexts such as exoplanetary atmospheres, cool stars and the interstellar medium. However, there are only very few astrophysical records of BeH, for example, a detection of $A^2\Sigma^+ \rightarrow X^2\Sigma^+$ emission lines of BeH in the sunspot umbra spectra by Wöhl (1971) and Shanmugavel et al. (2008).



Contents lists available at ScienceDirect

Atomic Data and Nuclear Data Tables

journal homepage: www.elsevier.com/locate/adt

Low-energy collisions between electrons and BeH^+ : Cross sections and rate coefficients for all the vibrational states of the ion

S. Niyonzima^{a,b}, S. Ilie^{a,c}, N. Pop^{a,c}, J. Zs. Mezei^{a,d,e,f}, K. Chakrabarti^g, V. Morel^h, B. Peres^h, D.A. Littleⁱ, K. Hassouni^d, Å. Larson^j, A.E. Orel^k, D. Benredjem^e, A. Bultel^h, J. Tennysonⁱ, D. Reiter^l, I.F. Schneider^{a,e,*}

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^f Institute of Nuclear Research, Hungarian Academy of Sciences, Debrecen, Hungary

^g Department of Mathematics, Scottish Church College, Calcutta 700 006, India

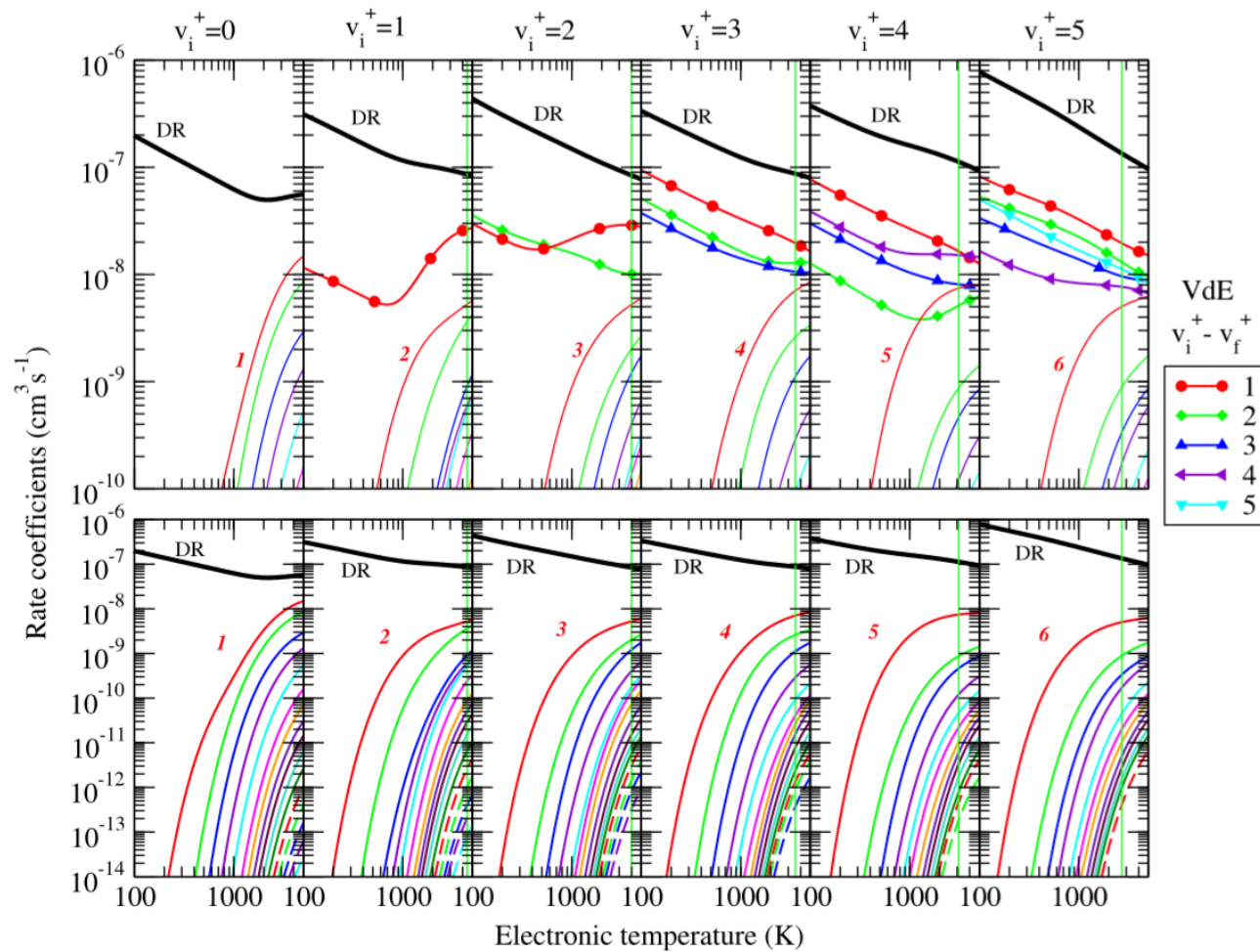
^h CORIA CNRS—Université de Rouen—Université Normandie, F-76801 Saint-Etienne du Rouvray, France

ⁱ Department of Physics and Astronomy, University College London, WC1E 6BT London, UK

^j Department of Physics, Stockholm University, AlbaNova University Center, S-106 91 Stockholm, Sweden

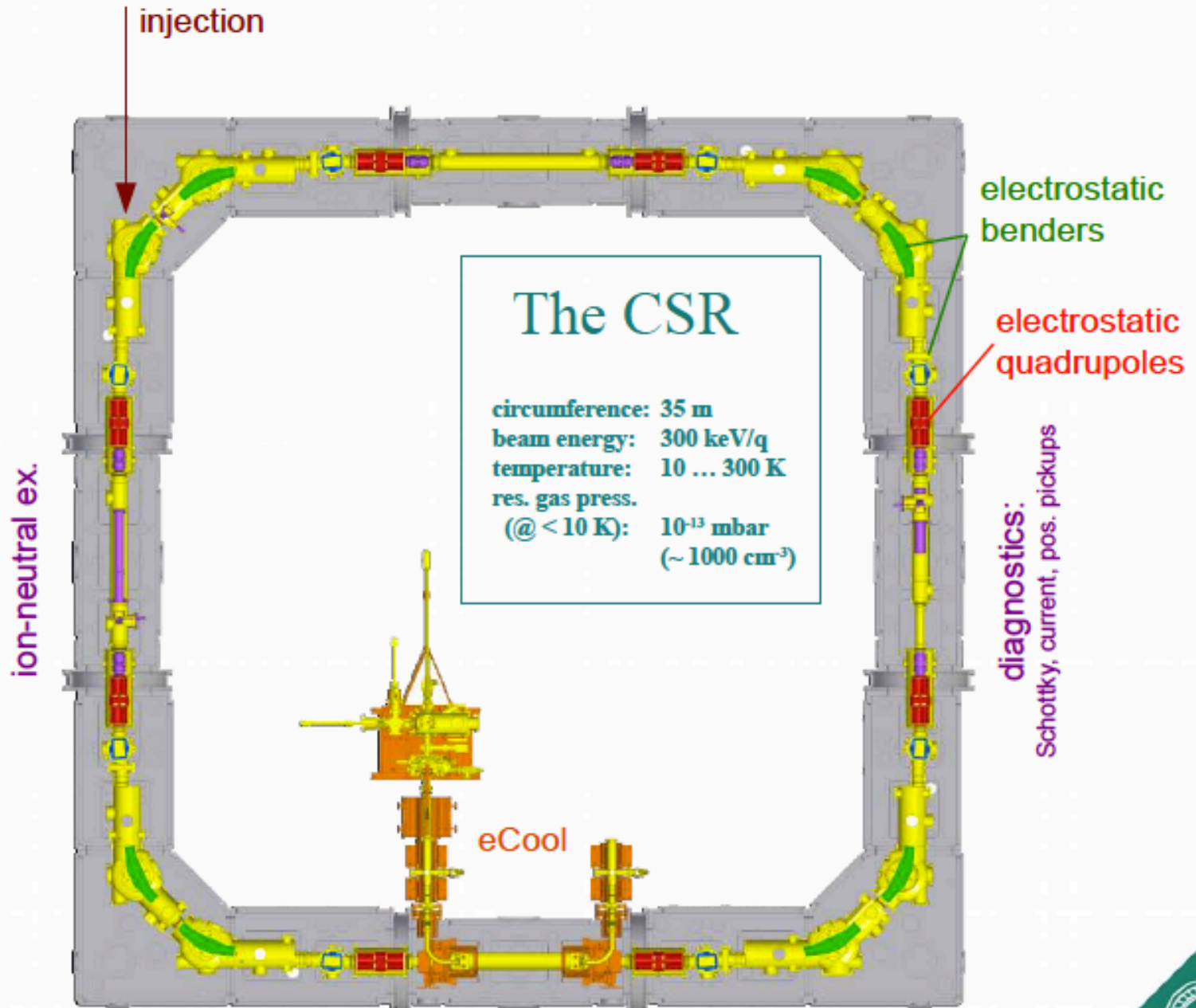
^k Department of Chemical Engineering and Materials Science, University of California, Davis, CA 95616, USA

^l Institute of Energy and Climate Research-Plasma Physics, Forschungszentrum Jülich GmbH Association EURATOM-FZJ, Partner in Trilateral Cluster, 52425 Jülich, Germany



Graph 4. Dissociative recombination (DR, thick line), vibrational excitation (VE, thin lines) and vibrational de-excitation (VdE, symbols and thick lines) Maxwell rate coefficients of ground ($v_i^+ = 0$) and excited ($v_i^+ = 1, \dots, 5$) BeH^+ in its electronic ground state (total mechanism). For VE, since the rate coefficients decrease monotonically with the excitation, the lowest final vibrational quantum number of the target is indicated only, and the lower panels extend the range down to $10^{-14} \text{ cm}^3/\text{s}$.

And ...
the FUTURE



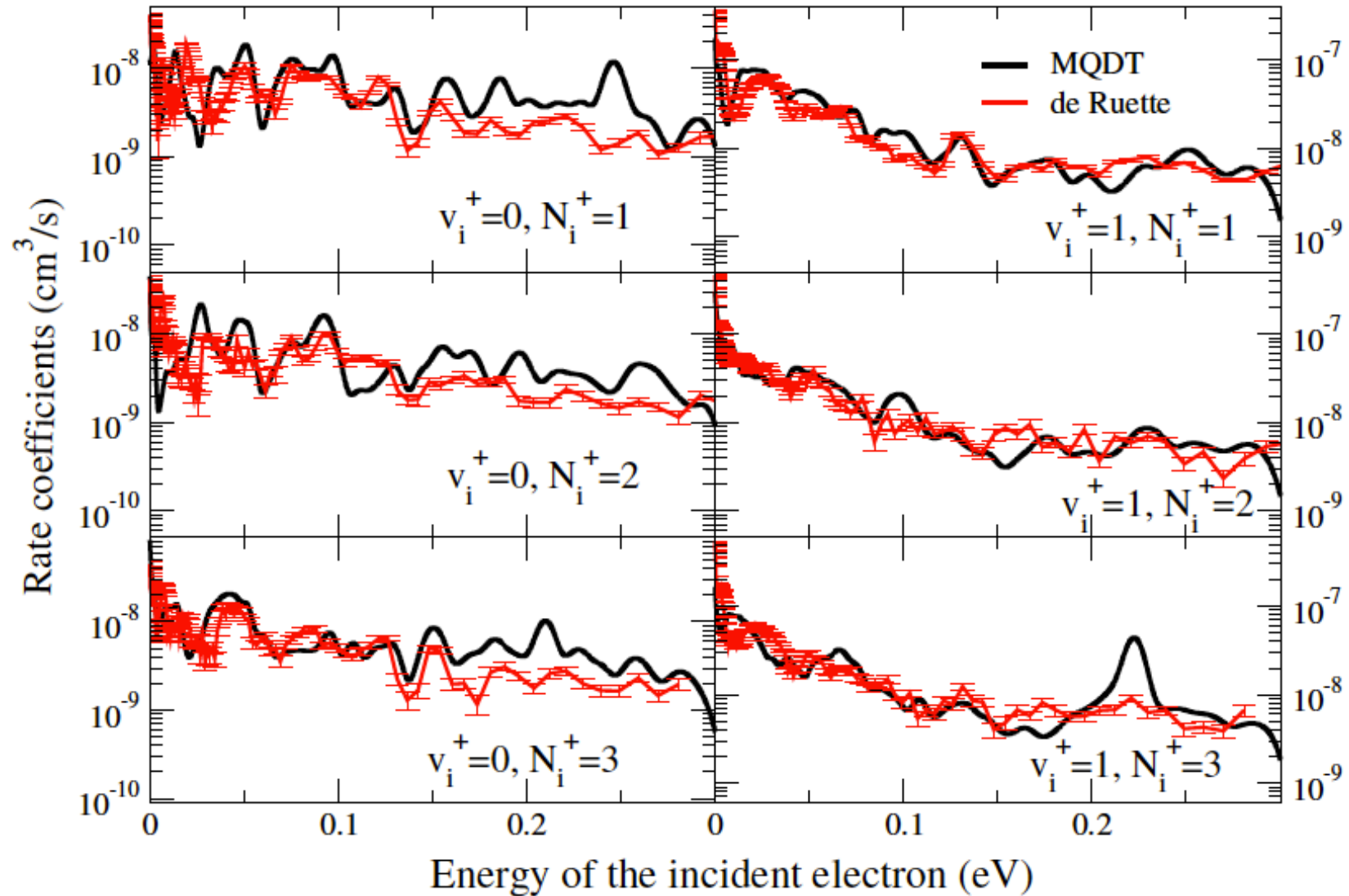
2016/12/12-Paris-HydrideToolbox



Are we prepared ?



1st state-to-state comparison experiment/theory ever !

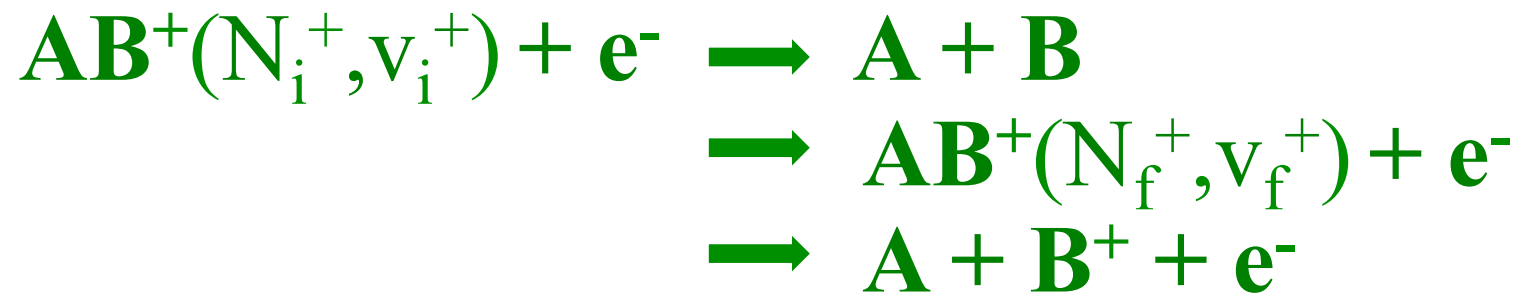


Perspectives:



...

**J. Zs Mezei¹⁻⁴, F. Colboc¹, D. O. Kashinsky⁵, D. A. Little⁶, M. D. Epée Epée⁷,
 S. Niyonzima⁸, N. Pop⁹, C. M. Coppola¹⁰, K. Chakrabarti¹¹, O. Motapon⁷,
 A. Bultel¹², K. Hassouni², D. Talbi¹³, A. P. Hickman¹⁴, A. Faure¹⁵, J. Tennyson⁶
 and I. F. Schneider^{1,3}**





**Thank you for
your attention !**