



## Chemistry of hydride-containing ices: UV photoprocessing of a H<sub>2</sub>O:CH<sub>4</sub> ice analog

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# 1. Introduction

## Astrophysical ice mantles composition

Low dust temperature

Molecules accrete  
on top of dust grains  
↓  
ice mantles

| Species                              | $X_{\text{H}_2\text{O}}$ [%]                 |                                                        |                                             |                                        |
|--------------------------------------|----------------------------------------------|--------------------------------------------------------|---------------------------------------------|----------------------------------------|
|                                      | MYSOs                                        | LYSOs                                                  | BG Stars                                    | Comets                                 |
| <b>Securely identified species :</b> |                                              |                                                        |                                             |                                        |
| H <sub>2</sub> O                     | 100                                          | 100                                                    | 100                                         | 100                                    |
| CO                                   | 7 <sub>4</sub> <sup>15</sup> (7)<br>3-26     | 21 <sub>12</sub> <sup>35</sup> (18)<br>(<3)-85         | 25 <sub>20</sub> <sup>43</sup><br>9-67      | 0.4-30                                 |
| CO <sub>2</sub>                      | 19 <sub>12</sub> <sup>25</sup><br>11-27      | 28 <sub>23</sub> <sup>37</sup><br>12-50                | 26 <sub>18</sub> <sup>39</sup><br>14-43     | 15 <sub>10</sub> <sup>24</sup><br>4-30 |
| CH <sub>3</sub> OH                   | 9 <sub>5</sub> <sup>23</sup> (5)<br>(< 3)-31 | 6 <sub>5</sub> <sup>12</sup> (5)<br>(< 1)-25           | 8 <sub>6</sub> <sup>10</sup> (6)<br>(<1)-12 | 0.2-7                                  |
| NH <sub>3</sub>                      | ~7 <sup>f</sup>                              | 6 <sub>4</sub> <sup>8</sup> (4)<br>3-10                | < 7                                         | 0.2-1.4                                |
| CH <sub>4</sub>                      | 1-3                                          | 4.5 <sub>3</sub> <sup>6</sup> (3)<br>1-11              | < 3                                         | 0.4-1.6                                |
| <b>Likely identified species :</b>   |                                              |                                                        |                                             |                                        |
| H <sub>2</sub> CO                    | ~2-7                                         | ~6                                                     |                                             | 0.11-1.0                               |
| OCN <sup>-</sup>                     | 0.6 <sub>0.3</sub> <sup>0.7</sup><br>0.1-1.9 | 0.6 <sub>0.4</sub> <sup>0.8</sup> (0.4)<br>(< 0.1)-1.1 | <0.5                                        |                                        |
| OCS                                  | 0.03-0.16                                    | ≤1.6                                                   | <0.22                                       | 0.1-0.4                                |

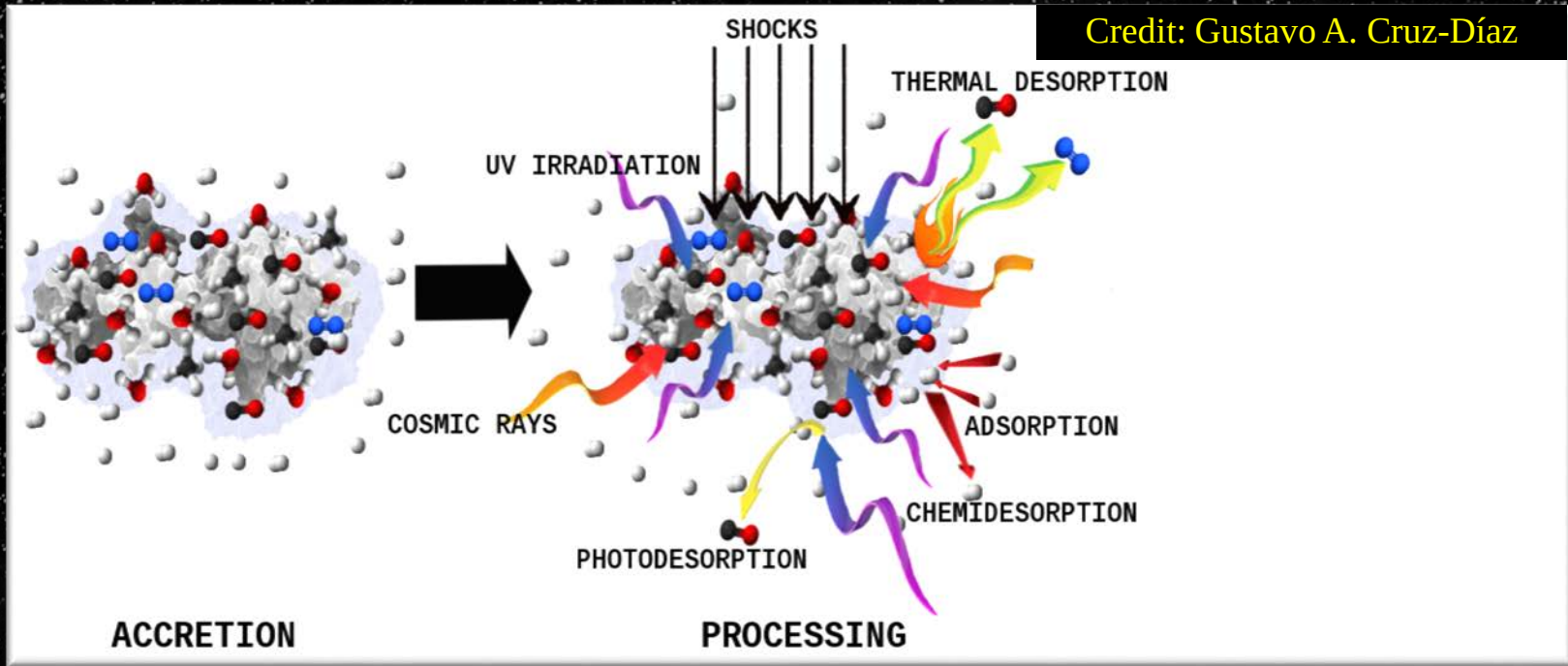
More species possibly present (e.g., N<sub>2</sub>, O<sub>2</sub>)

Boogert et al. 2015



# 1. Introduction

## Energetic processing of ice mantles



Ice mantles are energetically processed in dense cloud interiors

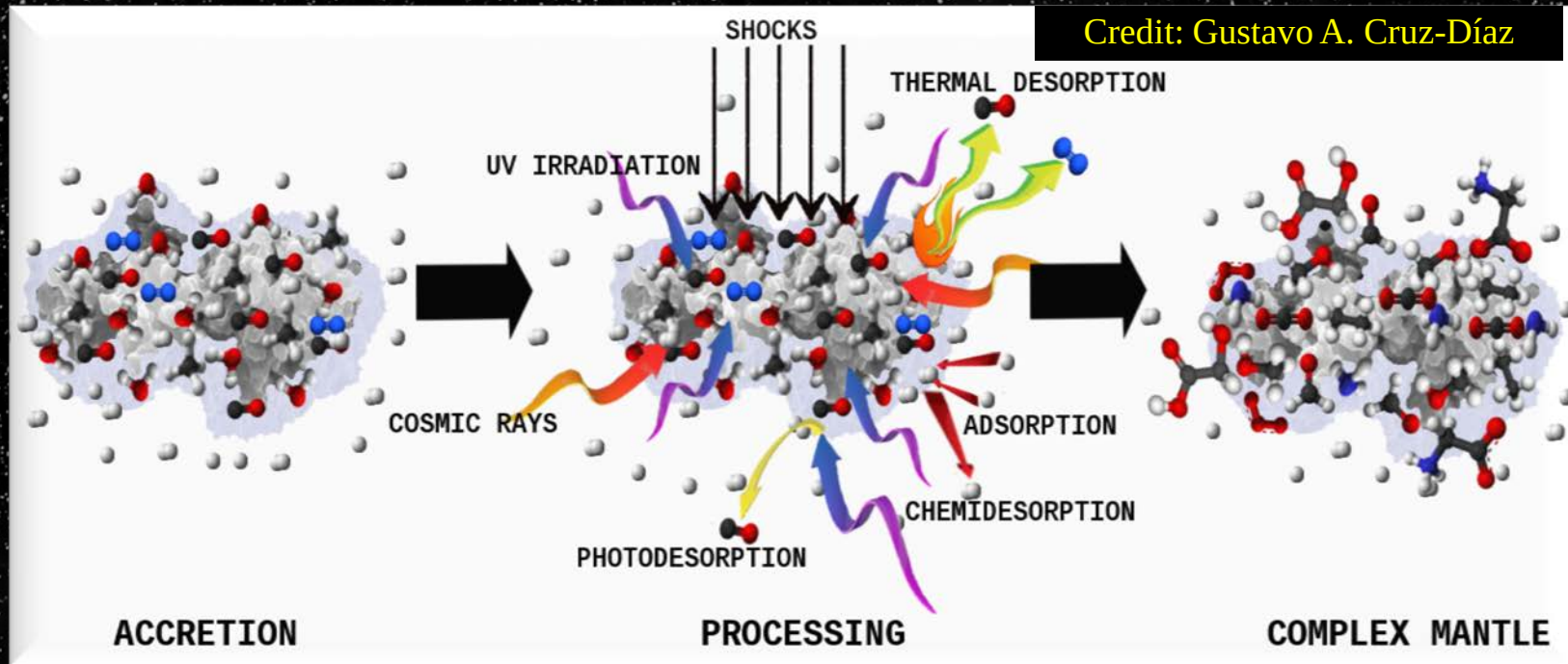
- Shocks
- Cosmic rays
- UV photons and X-rays

in circumstellar environments

- Thermal processing

# 1. Introduction

## Energetic processing of ice mantles



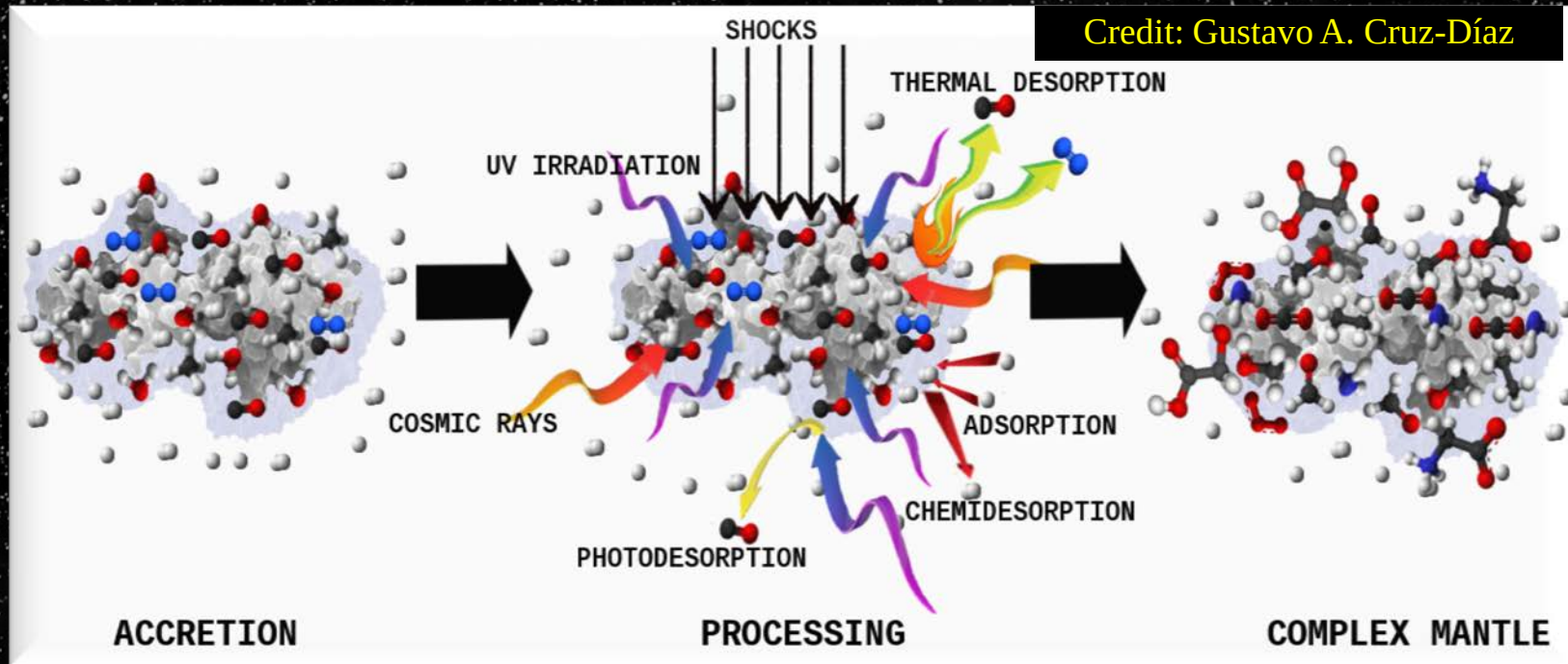
Ice mantles are energetically processed in dense cloud interiors

- Shocks → Sputtering
- Cosmic rays → Non-thermal desorption / ice chemistry
- UV photons and X-rays → Non-thermal desorption / ice chemistry  
in circumstellar environments
- Thermal processing → Thermal desorption /ice chemistry



# 1. Introduction

## Energetic processing of ice mantles



Ice mantles are energetically processed in dense cloud interiors

- Shocks → Sputtering
- Cosmic rays → Non-thermal desorption / ice chemistry
- **UV photons** and X-rays → **Non-thermal desorption / ice chemistry**  
in circumstellar environments
- Thermal processing → Thermal desorption /ice chemistry



## 2. Experimental set-up

### InterStellar Astrochemistry Chamber (ISAC)

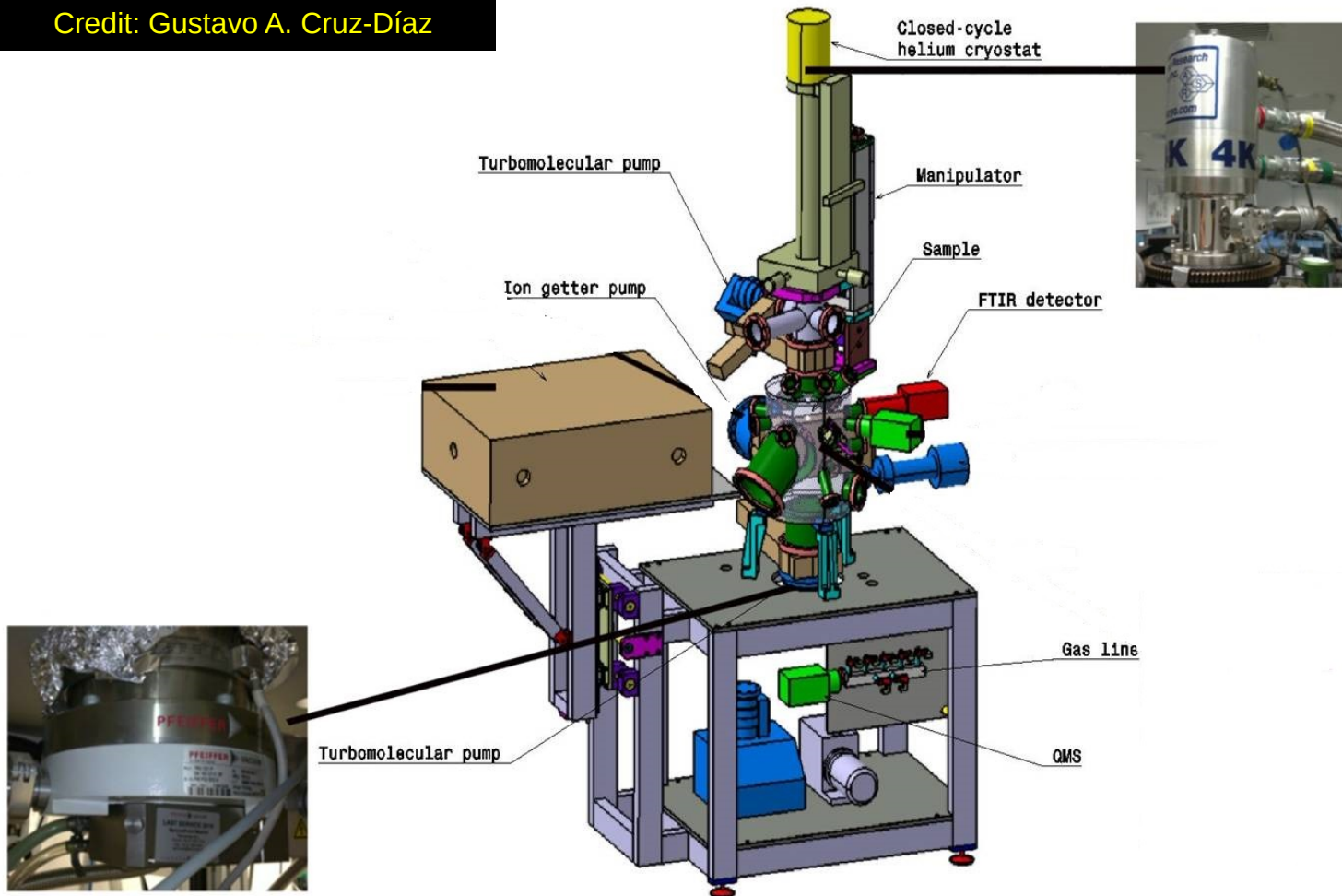


ISAC is an UHV chamber with  $P \sim 4 \cdot 10^{-11}$  mbar (dense clouds interiors)

## 2. Experimental set-up

### InterStellar Astrochemistry Chamber (ISAC)

Credit: Gustavo A. Cruz-Díaz

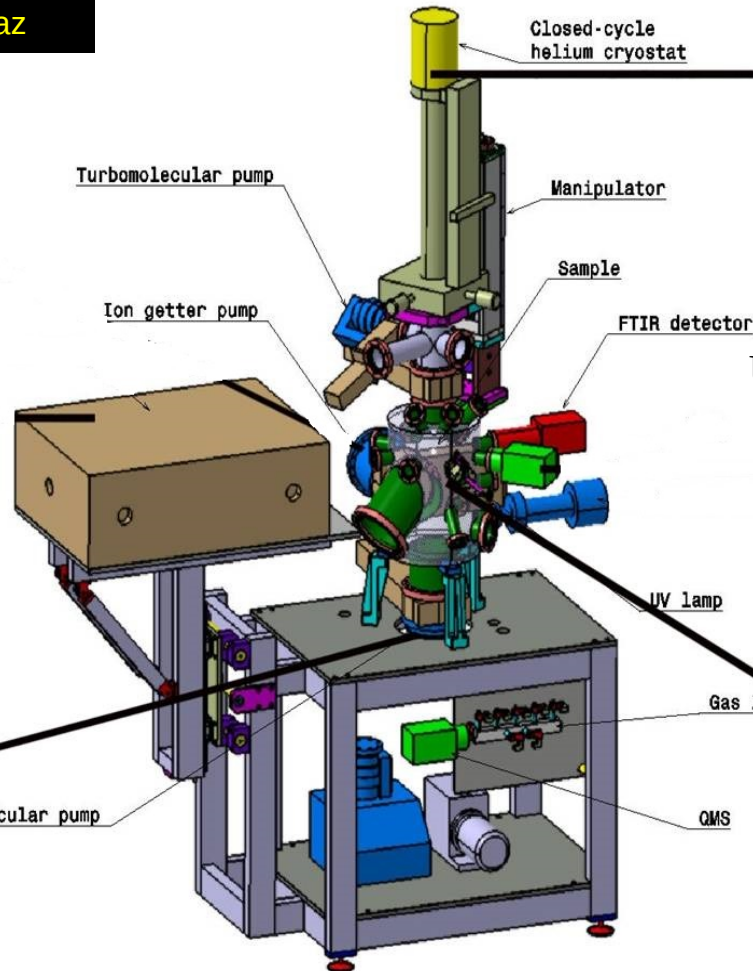




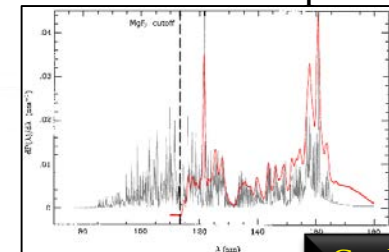
## 2. Experimental set-up

### InterStellar Astrochemistry Chamber (ISAC)

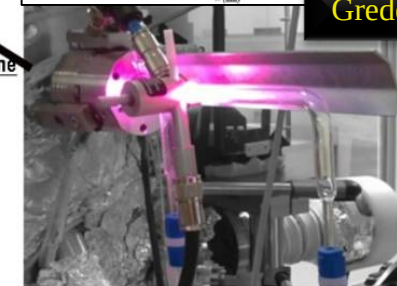
Credit: Gustavo A. Cruz-Díaz



UV flux =  $2 \times 10^{14}$  ph./cm<sup>2</sup> s



Gredel et al. 1989

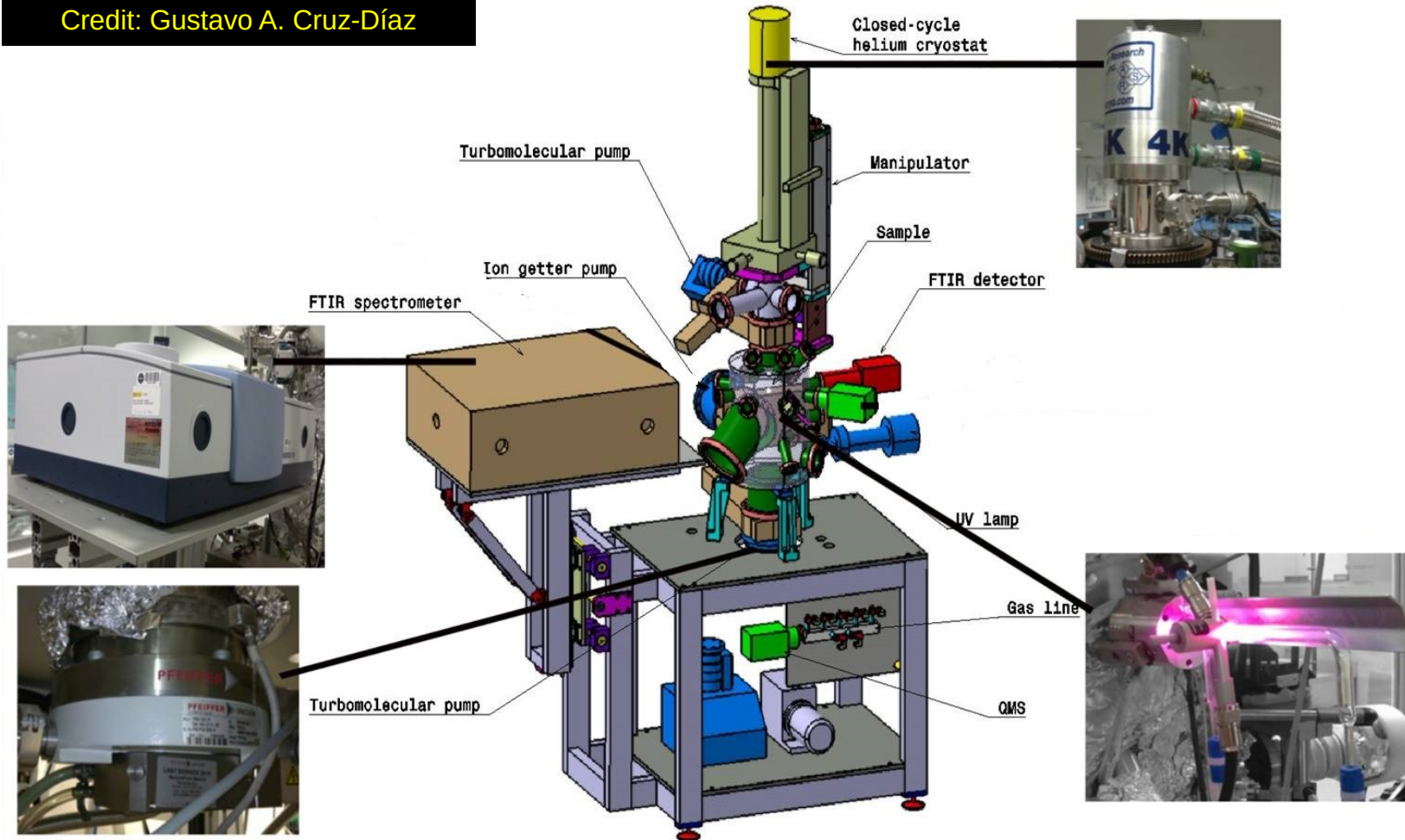




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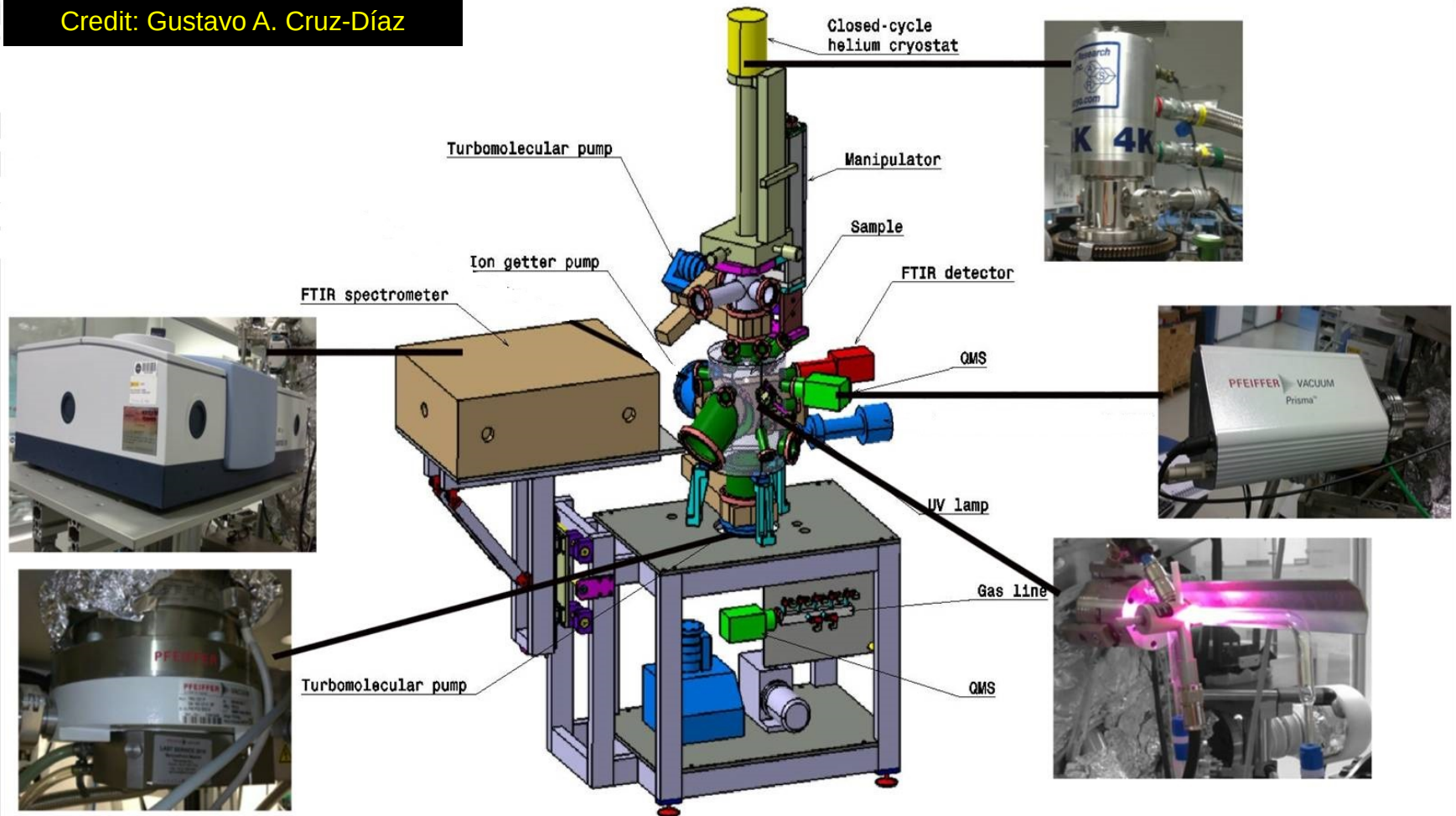
Credit: Gustavo A. Cruz-Díaz



## 2. Experimental set-up

### InterStellar Astrochemistry Chamber (ISAC)

Credit: Gustavo A. Cruz-Díaz







### 3. Photodesorption mechanisms

Ice molecules divided in 2 groups depending on  $E_d \leftrightarrow E_{hv}$

Different mechanisms?



### 3. Photodesorption mechanisms

Ice molecules divided in 2 groups depending on  $E_d \leftrightarrow E_{hv}$

#### *GROUP 1*

ices with (*almost*) NO photodissociation ( $E_{hv} < E_d$ )

Example: CO<sup>(1)</sup> ice, N<sub>2</sub> ice

<sup>(1)</sup>No direct photodissociation, only occurs via







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Example: CO<sup>(1)</sup> ice, N<sub>2</sub> ice

- 
- Photodesorption of original ice components
  - Photodesorption of photoproducts (no ph-chemistry)

<sup>(1)</sup>No direct photodissociation, only occurs via





### 3. Photodesorption mechanisms

Ice molecules divided in 2 groups depending on  $E_d \leftrightarrow E_{hv}$

#### *GROUP 1*

Photodesorption of original ice components via:

- ~~Direct photodesorption~~ (see, e.g., Bertin et al. 2012)
- Indirect photodesorption (DIET)

(Muñoz Caro et al. 2010, Fayolle et al. 2011, 2013, Bertin et al. 2012, 2014, Chen et al. 2014)





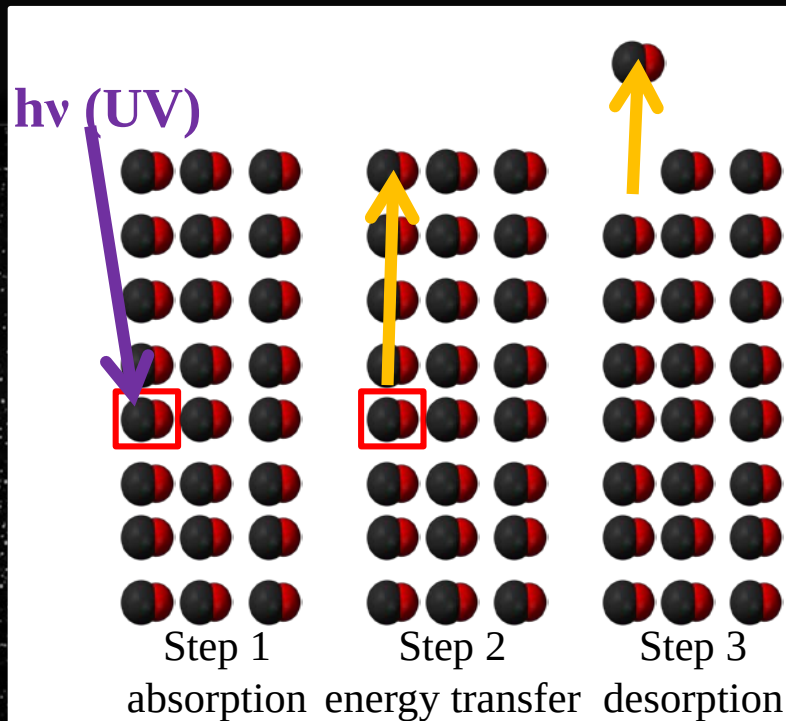
### 3. Photodesorption mechanisms

Ice molecules divided in 2 groups depending on  $E_d \leftrightarrow E_{hv}$

#### GROUP 1

Photodesorption of original ice components via:

- Direct desorption (see, e.g., Bertin et al. 2012)
- Indirect desorption (DIET)



al. 2011, 2013, Bertin et al. 2012, 2014,

Indirect desorption induced by electronic transitions (**DIET**): Electronic excitation energy is transferred to neighbors and provides energy to surface molecules breaking intermolecular bonds.



### 3. Photodesorption mechanisms

Ice molecules divided in 2 groups depending on  $E_d \leftrightarrow E_{hv}$

#### *GROUP 2*

ices with photodissociation ( $E_{hv} < E_d$ )

Example:  $\text{CO}_2$  ice,  $\text{CH}_3\text{OH}$  ice,  $\text{CH}_4$  ice, etc. (most molecules)





### 3. Photodesorption mechanisms

Ice molecules divided in 2 groups depending on  $E_d \leftrightarrow E_{hv}$

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  - Photodesorption of photoproducts (photochemistry)



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Ice molecules divided in 2 groups depending on  $E_d \leftrightarrow E_{hv}$

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- Photodesorption of original ice components
  - Photodesorption of photoproducts (photochemistry)



Additional photodesorption mechanisms!!



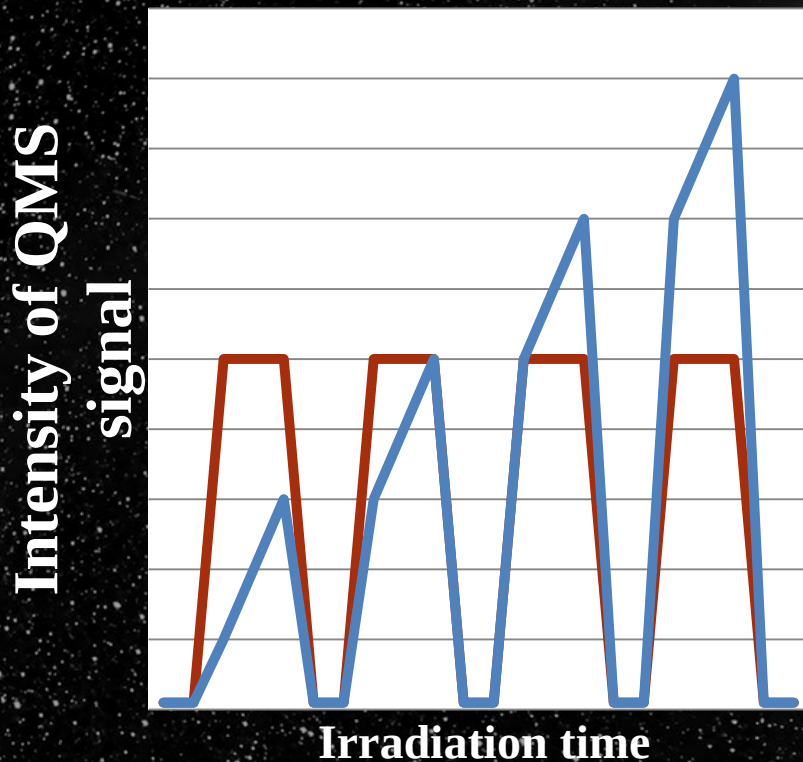


### 3. Photodesorption mechanisms

## Additional desorption mechanisms upon photodissociation

### GROUP 2

Photoproducts desorb with two yield vs fluence patterns



Two different photodesorption patterns

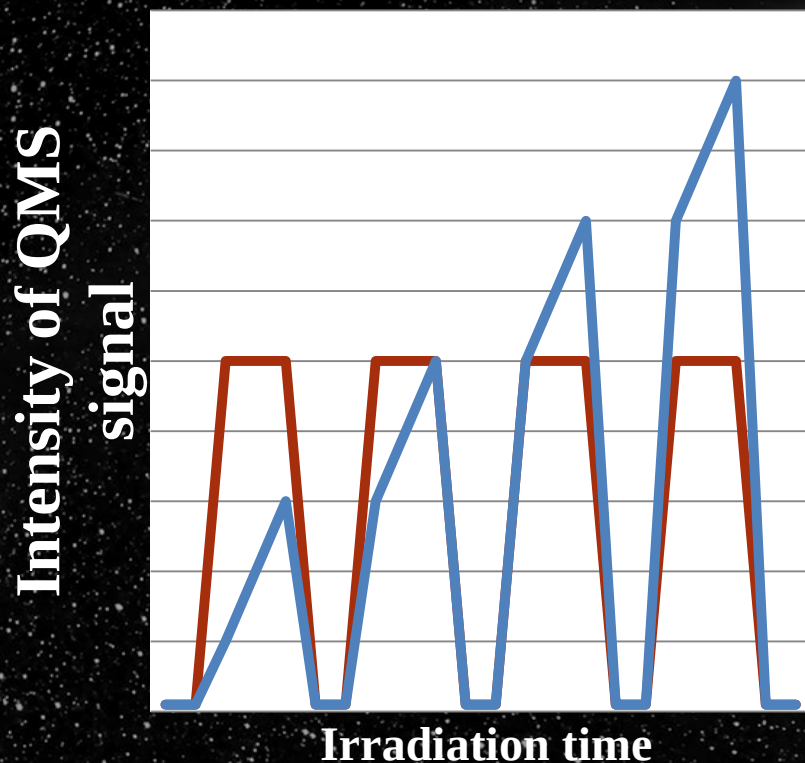


Two different photodesorption mechanisms

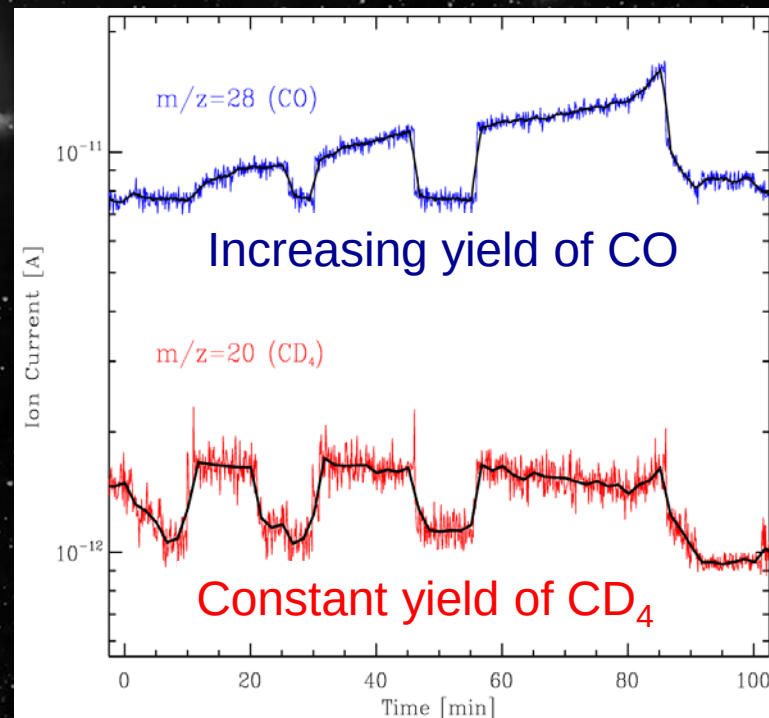
## Additional desorption mechanisms upon photodissociation

### GROUP 2

Photoproducts desorb with two  
yield vs fluence patterns



CD<sub>3</sub>OD ice irradiation







### 3. Photodesorption mechanisms

Increasing yield  $\leftrightarrow$  indirect desorption of photoproducts (DIET)

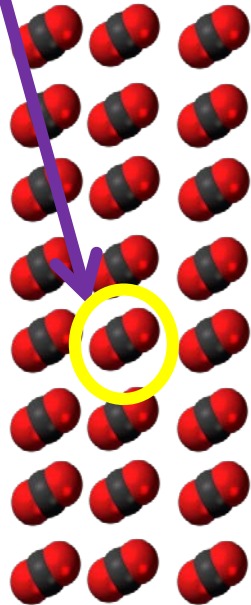


### 3. Photodesorption mechanisms

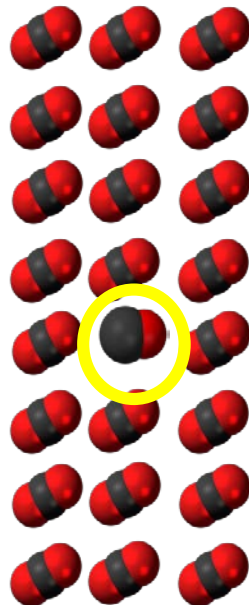
Increasing yield  $\leftrightarrow$  indirect desorption of photoproducts (DIET)

**Photoproducts formed in the ice bulk (first photon)**

$h\nu$  (first photon)



Step 1  
absorption



Step 2  
formation



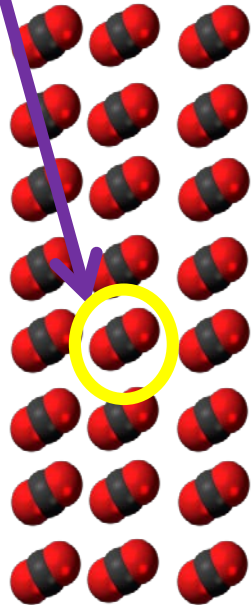


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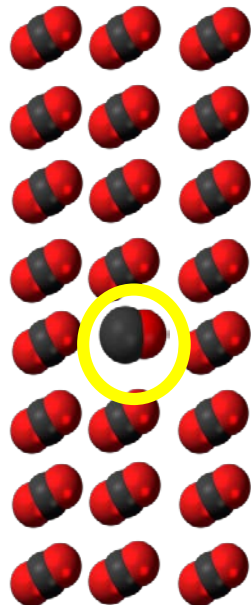
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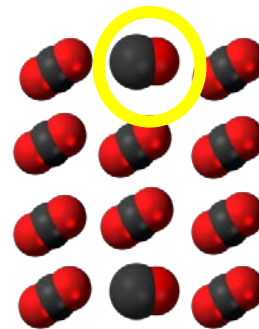
$h\nu$  (first photon)



Step 1  
absorption



Step 2  
formation



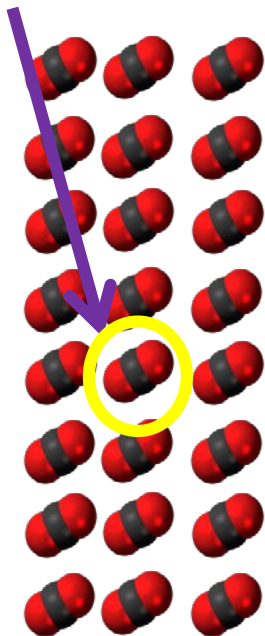


### 3. Photodesorption mechanisms

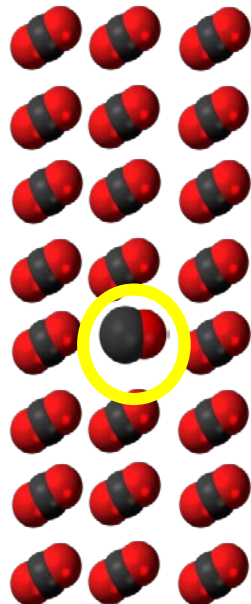
Increasing yield  $\leftrightarrow$  indirect desorption of photoproducts (DIET)

Photoproducts formed in the ice bulk (first photon)  
can desorb **LATER** via DIET mechanism (second photon)

$h\nu$  (first photon)

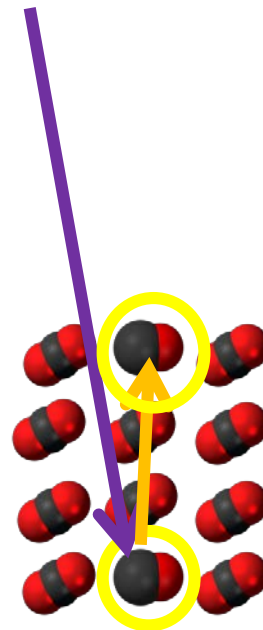


Step 1  
absorption



Step 2  
formation

$h\nu$  (second photon)



Steps 4 & 5  
absorption & energy transfer



Step 6  
desorption



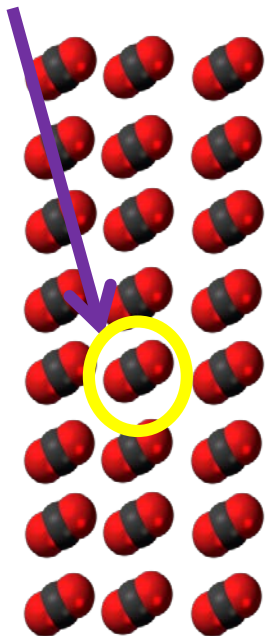


### 3. Photodesorption mechanisms

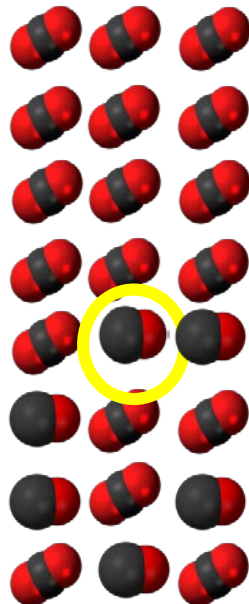
Increasing yield  $\leftrightarrow$  indirect desorption of photoproducts (DIET)

Photoproduct **ACCUMULATES** in the ice during irradiation  
 $\rightarrow$  photodesorption yield **INCREASES** with fluence

$h\nu$  (first photon)

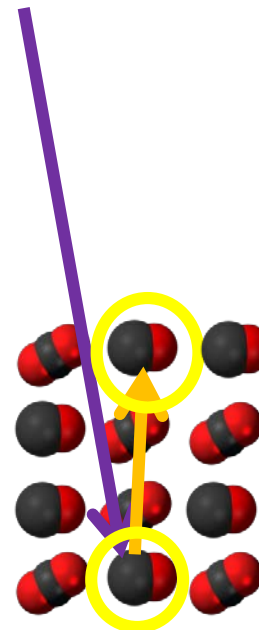


Step 1  
absorption

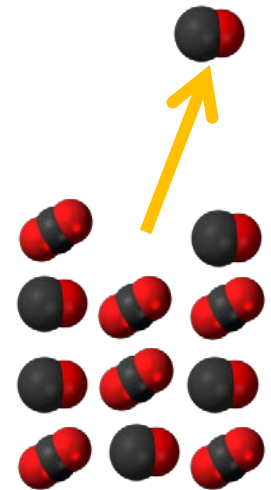


Step 2  
formation

$h\nu$  (second photon)



Steps 4 & 5  
absorption & energy transfer



Step 6  
desorption



### 3. Photodesorption mechanisms

Additional desorption mechanism upon photodissociation

Constant yield  $\leftrightarrow$  immediate desorption of photoproducts





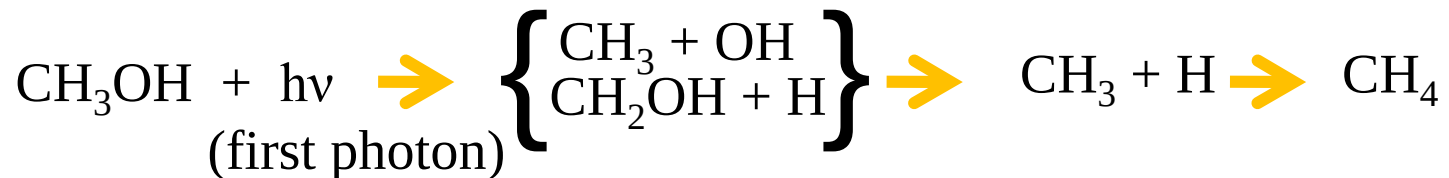
### 3. Photodesorption mechanisms

## Additional desorption mechanism upon photodissociation

Constant yield  $\leftrightarrow$  immediate desorption of photoproducts

**Photoproducts formed on the ice surface (first photon)**

- by photofragmentation
- by recombination of radicals



$\text{CH}_3\text{OH}$  ice surface

Cruz-Díaz et al. 2016, A&A 592, A68



### 3. Photodesorption mechanisms

## Additional desorption mechanism upon photodissociation

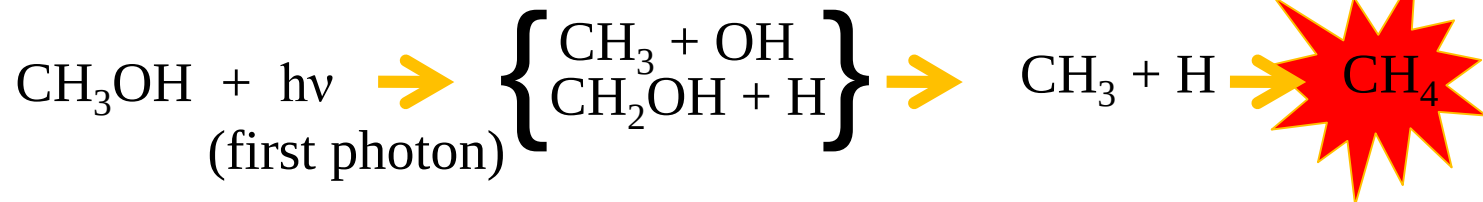
Constant yield  $\leftrightarrow$  immediate desorption of photoproducts

**Photoproducts formed on the ice surface (first photon)**

- by photofragmentation
- by recombination of radicals

can desorb **IMMEDIATELY** because of the excess energy (no more photons needed)

Fayolle et al. 2013, Fillion et al. 2014



$\text{CH}_3\text{OH}$  ice surface

Cruz-Díaz et al. 2016, A&A 592, A68



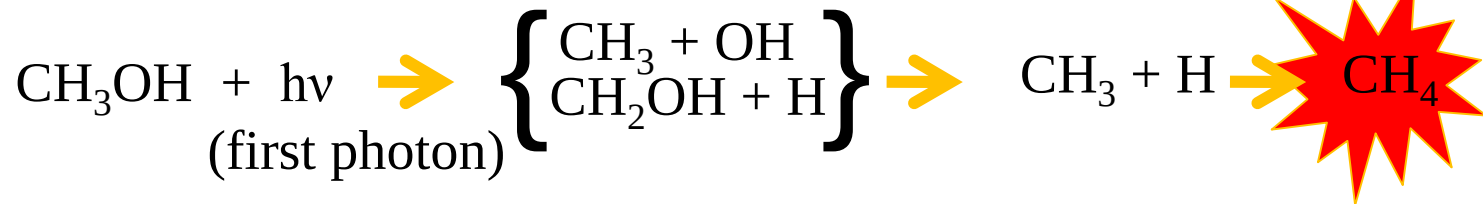


### 3. Photodesorption mechanisms

## Additional desorption mechanism upon photodissociation

Constant yield  $\leftrightarrow$  immediate desorption of photoproducts

**Photoproducts DO NOT ACCUMULATE on the ice surface during irradiation**  
 **$\rightarrow$  photodesorption yield is CONSTANT with fluence**



$\text{CH}_3\text{OH}$  ice surface



### 3. Photodesorption mechanisms

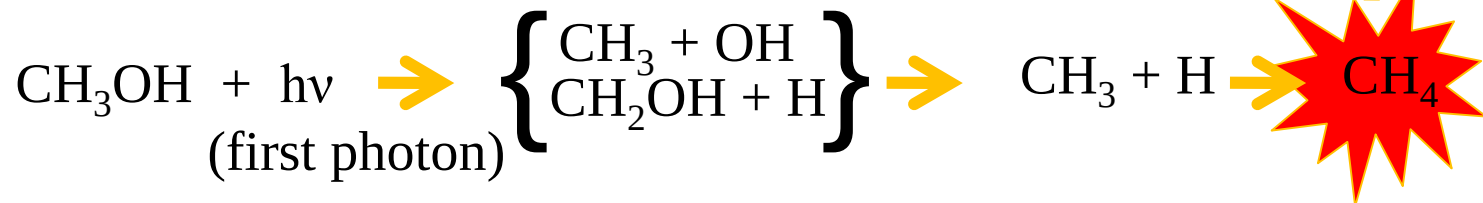
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Constant yield  $\leftrightarrow$  immediate desorption of photoproducts

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### PHOTOCHEMIDESORPTION

Martín-Doménech et al. 2016, A&A 589, A107  
Cruz Díaz et al. 2016, A&A 592, A68



$\text{CH}_3\text{OH}$  ice surface

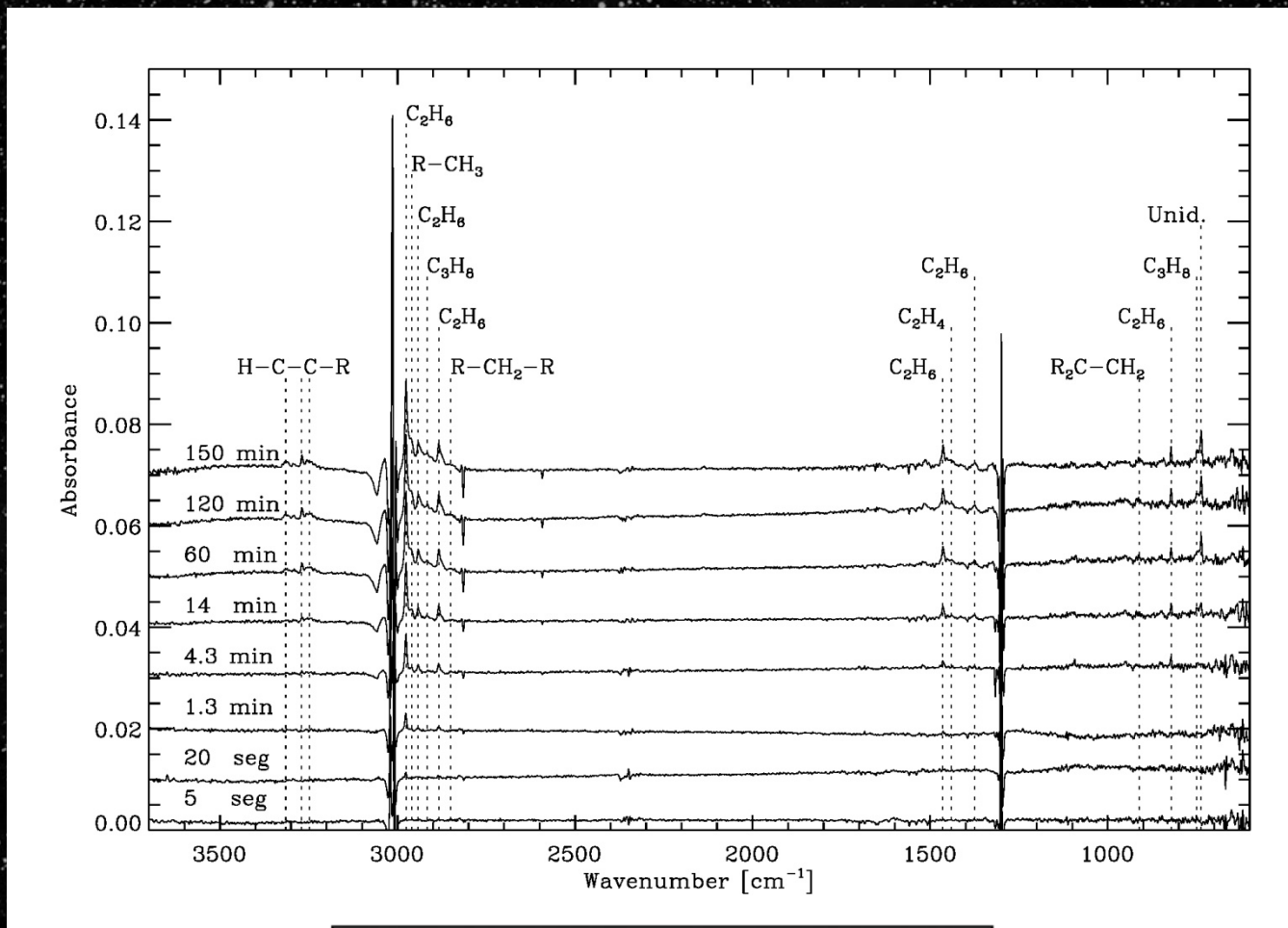
Cruz-Díaz et al. 2016, A&A 592, A68





## 4. Photoprocessing of CH<sub>4</sub> ice

### Photodissociation of CH<sub>4</sub> → Formation of photoproducts

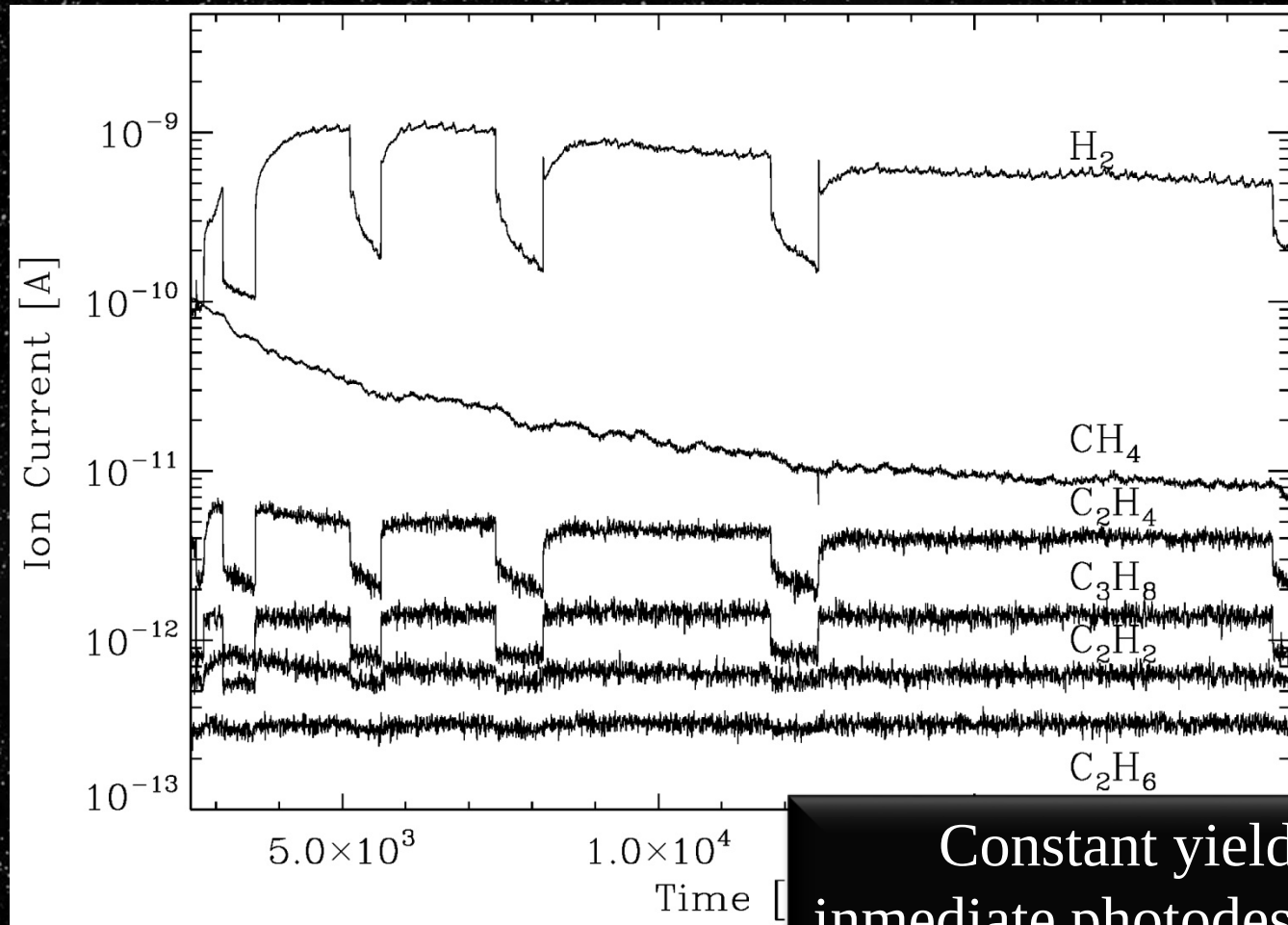


Cruz-Díaz et al., *in prep.*



## 4. Photoprocessing of CH<sub>4</sub> ice

### Photodesorption of photoproducts observed by QMS



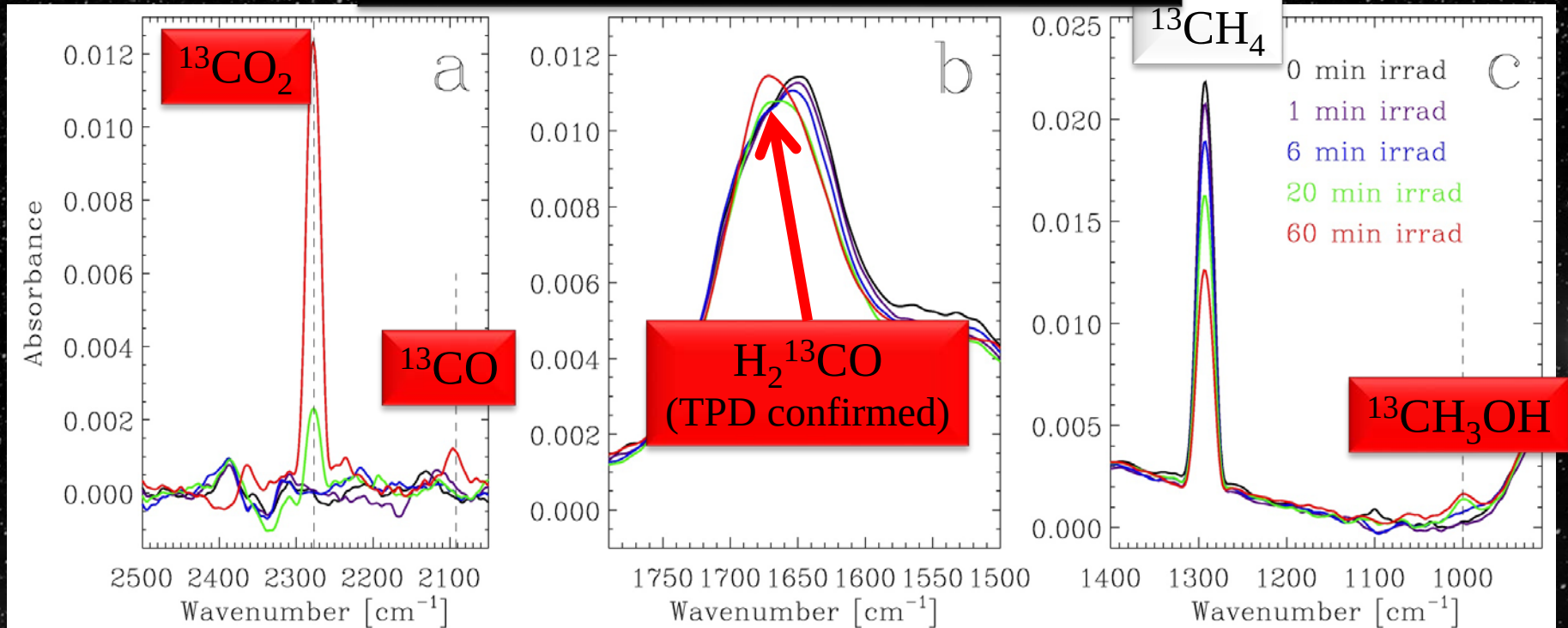
Cruz-Díaz et al., *in press*

Constant yield =  
immediate photodesorption  
(photochemidesorption)



Formation of photoproducts  $\rightarrow$  new IR bands during irradiation

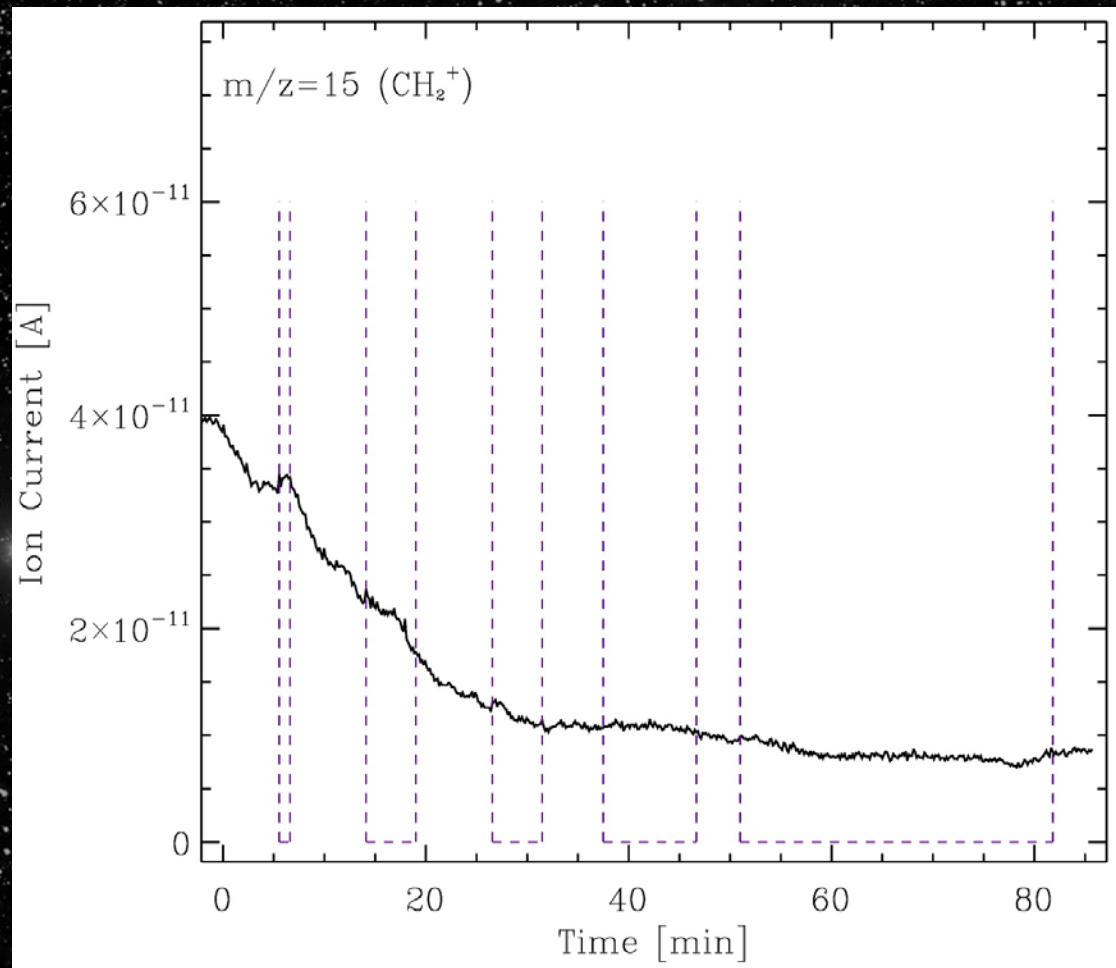
Martín-Doménech et al. 2016, A&amp;A 589, A107





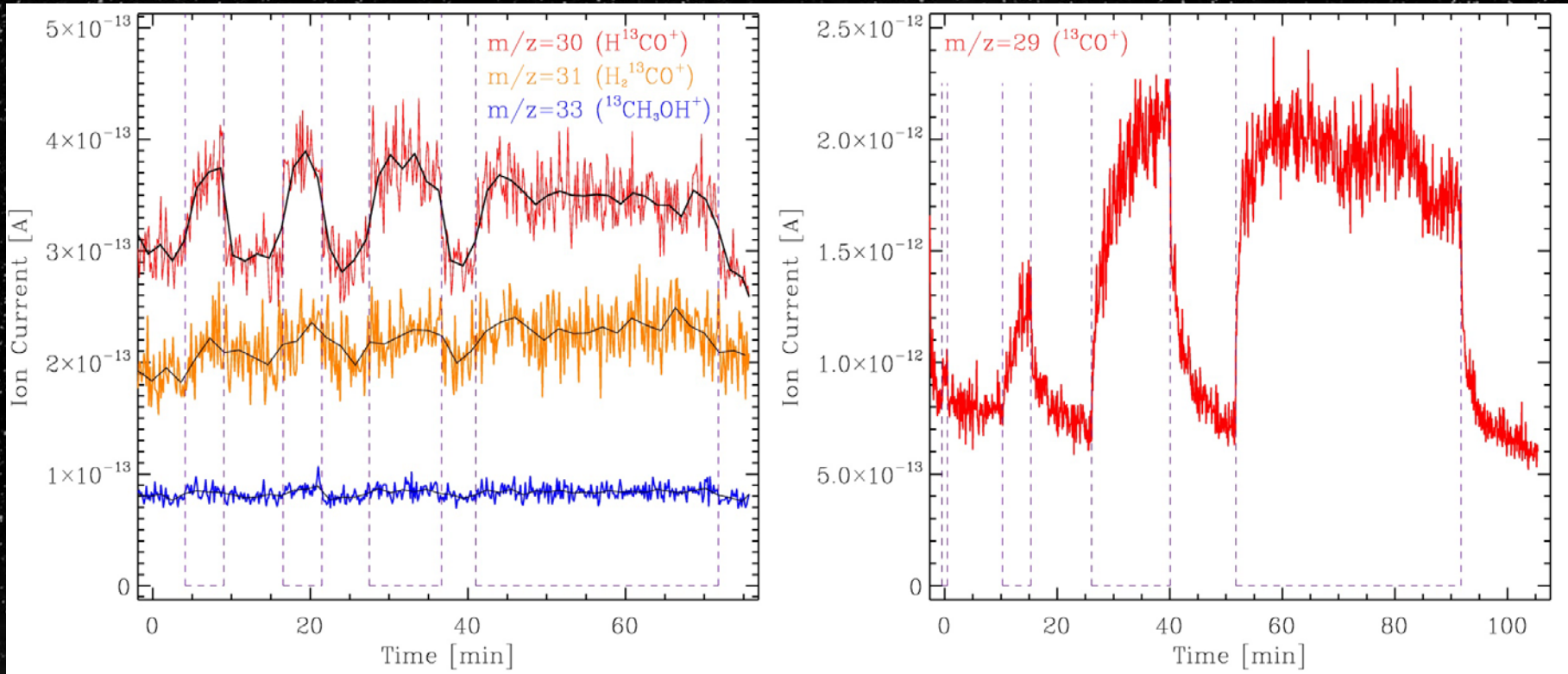
## 5. Photoprocessing of $\text{H}_2\text{O}:\text{}^{13}\text{CH}_4$

### Photodesorption of $\text{CH}_4$ not observed by QMS





# Photodesorption of $\text{H}_2\text{CO}$ and $\text{CO}$ observed by QMS

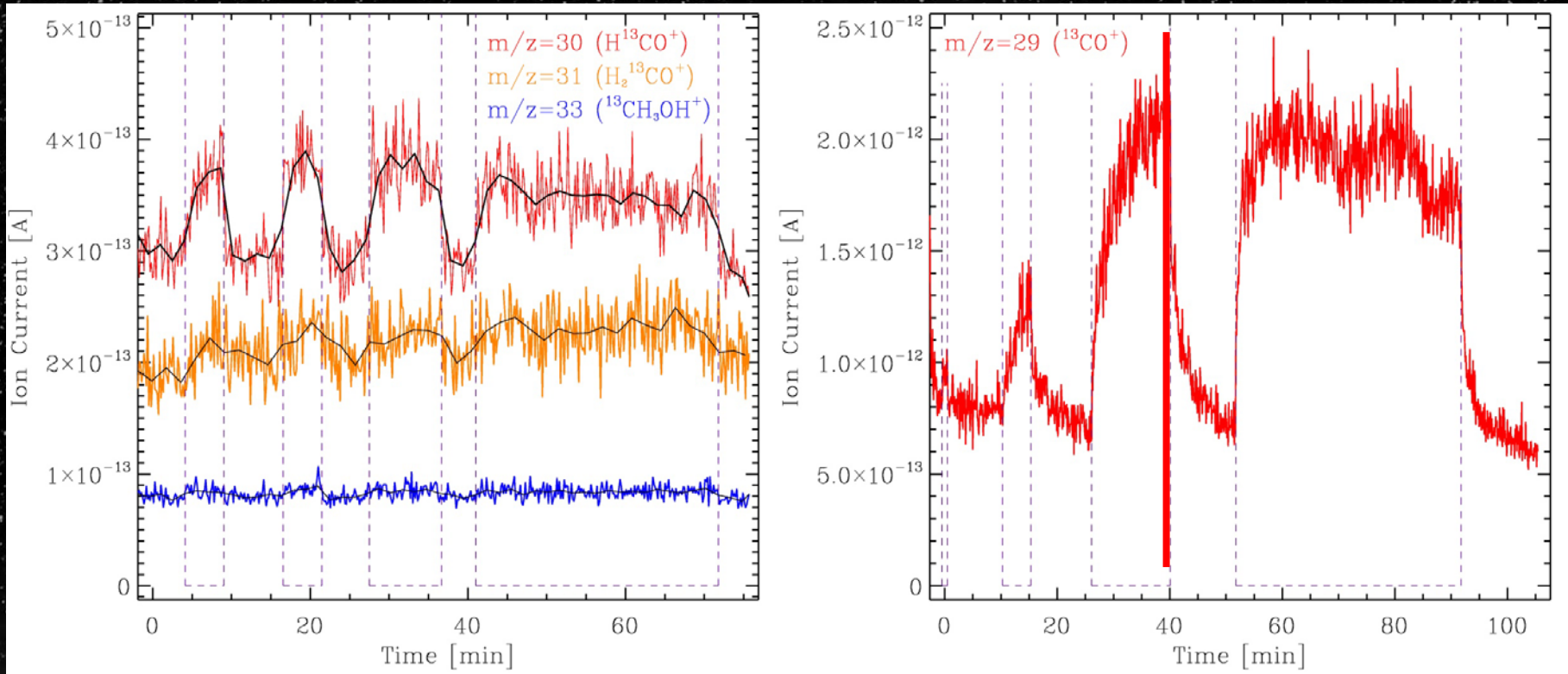


Martín-Doménech et al. 2016, A&A 589, A107

Constant yield =  
immediate photodesorption  
(photochemidesorption)

Increasing yield =  
indirect photodesorption  
(DIET)

# Photodesorption of $\text{H}_2\text{CO}$ and $\text{CO}$ quantified by QMS



Martín-Doménech et al. 2016, A&A 589, A107

$$Y_{\text{pd}}(\text{H}_2\text{CO}) = 4.4 \times 10^{-5} \text{ mol. / inc. ph.}$$

$$Y_{\text{pd}}(\text{CO}) = 1.5 \times 10^{-4} \text{ mol. / inc. ph.}$$





## Photodesorption of $\text{CH}_4$

$\text{CH}_4$  ice



$\text{H}_2\text{O}:\text{CH}_4$  ice

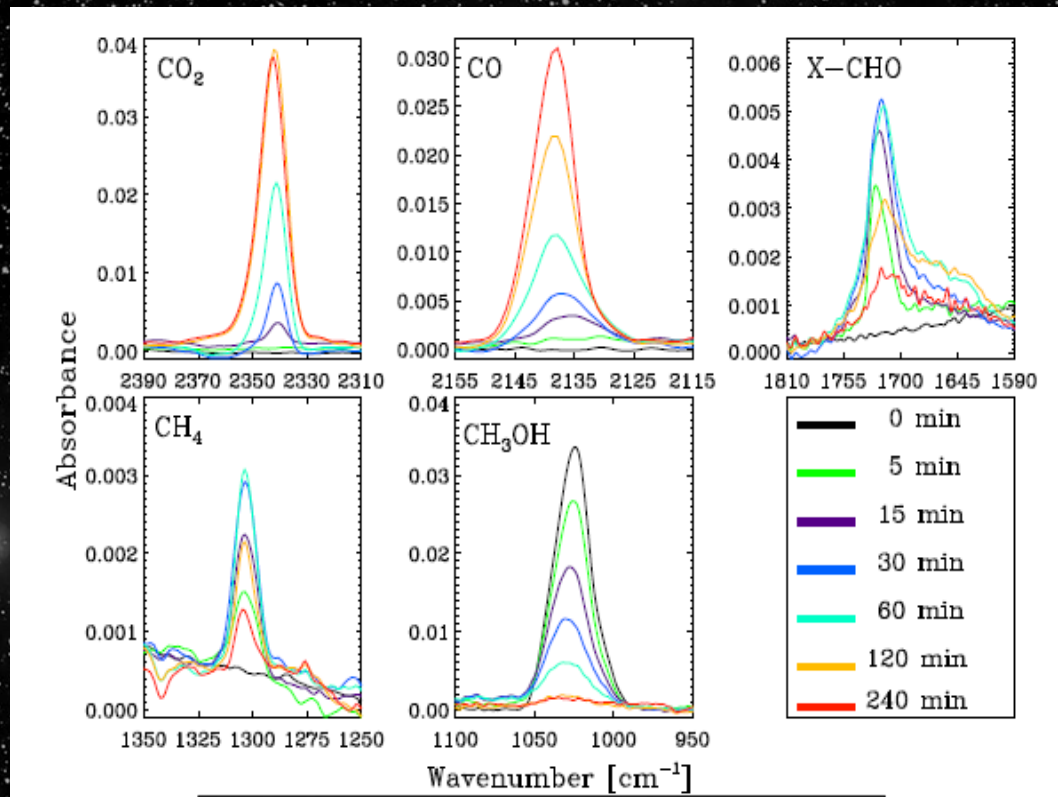


$\text{CH}_3\text{OH}$  ice



Additional desorption mechanism for photoproducts

## Formation of photoproducts → new IR bands during irradiation



Cruz-Díaz et al. 2016, A&amp;A 592, A68

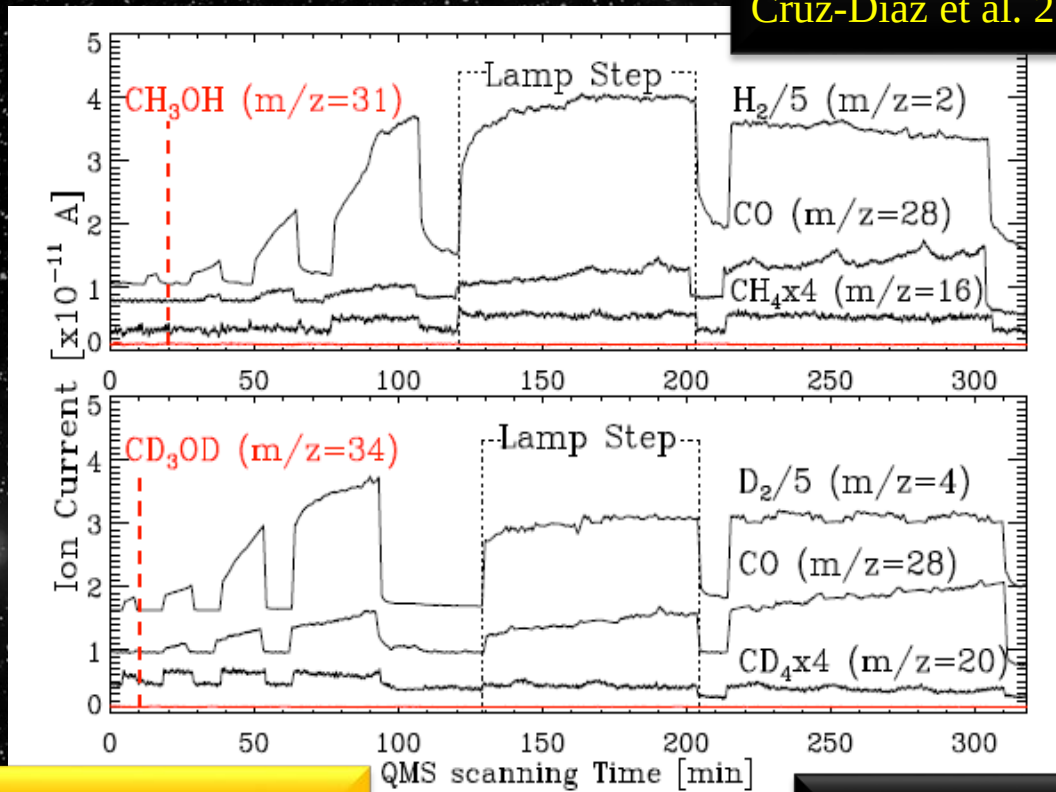




## 6. Photoprocessing of CH<sub>3</sub>OH ice

### Photodesorption of H<sub>2</sub>CO and CO observed by QMS

Cruz-Díaz et al. 2016, A&A 592, A68



CH<sub>4</sub> → Constant yield =  
immediate photodesorption  
(photochemidesorption)

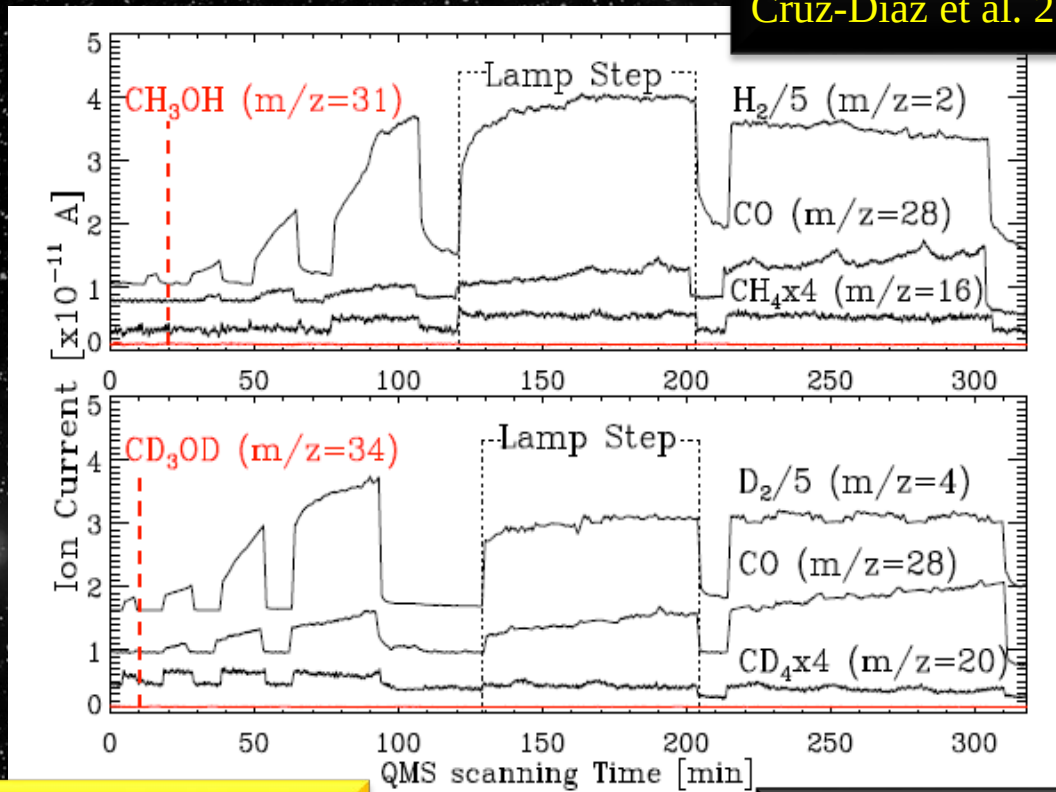
CO → Increasing yield =  
indirect photodesorption  
(DIET)



## 6. Photoprocessing of CH<sub>3</sub>OH ice

### Photodesorption of H<sub>2</sub>CO and CO observed by QMS

Cruz-Díaz et al. 2016, A&A 592, A68



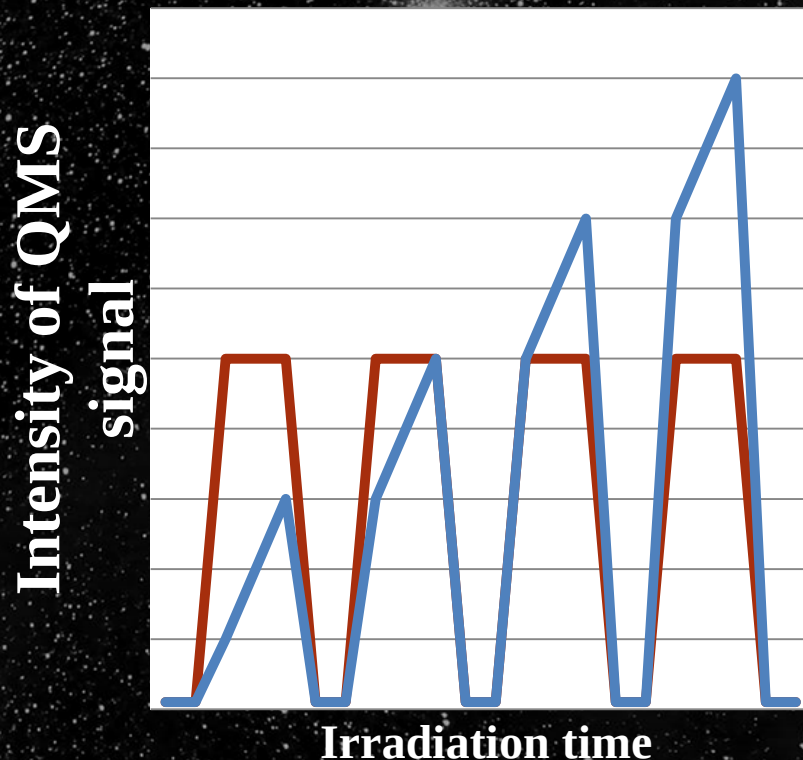
$$\text{CH}_4 \rightarrow Y_{\text{pd}} = 2.0 \times 10^{-3} \text{ mol. / inc. ph.}$$

$$\text{CO} \rightarrow Y_{\text{pd max}} = 9.0 \times 10^{-3} \text{ mol. / inc. ph.}$$



## 7. Conclusions

Photoproducts desorb with two  
yield vs fluence patterns



Increasing yield = INDIRECT  
desorption of accumulated  
photoproducts via ENERGY  
TRANSFER

Constant yield = IMMEDIATE  
desorption of formed  
photoproducts due to  
EXCESS ENERGY

### Photodesorption of CH<sub>4</sub>

CH<sub>4</sub> ice



H<sub>2</sub>O:CH<sub>4</sub> ice



CH<sub>3</sub>OH ice



Additional desorption mechanism for photoproducts





# Acknowledgments

## Colaborators:

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**Merci !**

**THE HYDRIDE TOOLBOX**

